Decay of standard-model-like Higgs boson $h \rightarrow \mu \tau$ in a 3-3-1 model with inverse seesaw neutrino masses

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By adding new gauge singlets of neutral leptons, the improved versions of the 3-3-1 models with right-handed neutrinos have been recently introduced in order to explain recent experimental neutrino oscillation data through the inverse seesaw mechanism. We prove that these models predict promising signals of lepton-flavor-violating decays of the standard-model-like Higgs boson $h_1^0 \to \mu \tau$, $e\tau$, which are suppressed in the original versions. One-loop contributions to these decay amplitudes are introduced in the unitary gauge. Based on a numerical investigation, we find that the branching ratios of the decays $h_1^0 \to \mu \tau$, $e\tau$ can reach values of 10^{-5} in the regions of parameter space satisfying the current experimental data of the decay $\mu \to e\gamma$. The value of 10^{-4} appears when the Yukawa couplings of leptons are close to the perturbative limit. Some interesting properties of these regions of parameter space are also discussed.

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I. INTRODUCTION

Signals of lepton-flavor-violating decays of the standard-model-like Higgs boson (LFVHDs) were investigated at the LHC [1] not very long after its discovery in 2012 [2]. So far, the most stringent limits on the branching ratios (Br) of these decays are $\text{Br}(h \to \mu \tau, e \tau) < \mathcal{O}(10^{-3})$ from the CMS Collaboration using data collected at a center-of-mass energy of 13 TeV. The sensitivities of the planned colliders for LFVHD searches are predicted to reach the order of 10^{-5} [3].

On the theoretical side, model-independent studies showed that the LFVHDs predicted from models beyond the standard model (BSM) are constrained indirectly from experimental data such as lepton-flavor-violating decays of charged leptons (cLFV) [4]. Namely, they are affected most strongly by the recent experimental bound on $\text{Br}(\mu \to e \gamma)$. Fortunately, large branching ratios of the decays $h \to \mu \tau, e \tau$

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³. are still allowed up to the order of 10^{-4} . Also, LFVHDs have been widely investigated in many specific BSM models, where the decay rates were indicated to be close to the upcoming sensitivities of colliders, including nonsupersymmetric [5,6] and supersymmetric versions [7]. Among them, the models based on the gauge symmetry $SU(3)_C \times SU(3)_L \times U(1)_X$ (3-3-1) contain rich lepton-flavor-violating (LFV) sources which may result in interesting cLFV phenomenology such as charged lepton decays $e_i \rightarrow e_j \gamma$ [8–11]. In particular, it was shown that $Br(\mu \rightarrow e \gamma)$ is large in these models, and hence it must be taken into account to constrain the parameter space. In addition, such rich LFV resources may give large LFVHD rates as promising signals of new physics.

Although the 3-3-1 models were introduced a long time ago [12,13], LFVHDs have been investigated only in the version with heavy neutral leptons assigned as the third components of lepton (anti)triplets, where active neutrino masses come from effective operators [14]. The largest values of LFVHD rates were shown to be $\mathcal{O}(10^{-5})$, originating from heavy neutrinos and charged Higgs bosons [6]. Improved versions consisting of new neutral lepton singlets were recently introduced [8,15]. They are more interesting because they successfully explain the experimental neutrino data through the inverse seesaw (ISS) mechanism. We call them the 331ISS models for short. They predict a large cLFV decay rate of $\mu \to e \gamma$

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corresponding to recent experimental bounds. They may also contain dark matter candidates [8,15]. These properties make them much more attractive than the original versions of 3-3-1 models with right-handed neutrinos (331RHN) [13]. They predict suppressed LFV decay rates, because all neutrinos including exotic ones are extremely light. Furthermore, loop corrections to the neutrino mass matrix must be taken into account to obtain an active neutrino mass spectrum that explains the experimental data [16]. Hence, LFV signals are an interesting way to distinguish the 331ISS and 331RHN models. More specifically, a simple ISS extension of the SM allows large $Br(h \to \mu \tau, e \tau) \sim$ $\mathcal{O}(10^{-5})$ in the allowed regions satisfying Br($\mu \to e\gamma$) < 4.2×10^{-13} [17]. Inspired by this, we will address the following questions in this work: how large is the Br($h \rightarrow \mu \tau, e \tau$) predicted by the 331ISS models under the experimental constraints of the cLFV decays?; and, are these branching ratios larger than the values calculated in the simplest ISS extension of the SM? Because these 331 models contain many more particles that contribute to LFV processes through loop corrections, either constructive or destructive correlations among them will strongly affect the allowed regions of the parameter space satisfying the current bound of the decay rate $\mu \to e\gamma$. The most interesting allowed regions will also allow large LFVHD rates, which we will try to look for in this work. Because the discussion on the decay $h \to e\tau$ is rather similar to the decay $h \to \mu\tau$, we only briefly mention the latter.

Our paper is organized as follows. In Sec. II we discuss the necessary ingredients of a 331ISS model for studying LFVHDs and how the ISS mechanism works to generate active neutrino parameters consistent with current experimental data. In Sec. III we present all couplings needed to determine the one-loop amplitudes of the LFVHDs of the SM-like Higgs boson. In Sec. IV we show important numerical LFVHD results predicted by the 331ISS model. Section V contains our conclusions. Finally, the Appendix lists all of the analytic formulas expressing one-loop contributions calculated in the unitary gauge.

II. THE 331ISS MODEL FOR TREE-LEVEL NEUTRINO MASSES

A. The model and neutrino masses from the inverse seesaw mechanism

First, we will consider a 331ISS model based on the original 331RHN model given in Ref. [16], where active neutrino masses and oscillations are generated from the ISS mechanism. The quark sector and $SU(3)_C$ representations are irrelevant in this work, and hence they are omitted here. The electric charge operator corresponding to the gauge group $SU(3)_L \times U(1)_X$ is $Q = T_3 - \frac{1}{\sqrt{3}}T_8 + X$, where $T_{3,8}$ are diagonal $SU(3)_L$ generators. Each lepton family consists of a $SU(3)_L$ triplet $\psi_{aL} = (\nu_a, e_a, N_a)_L^T \sim (3, -\frac{1}{3})$ and a right-handed charged lepton $e_{aR} \sim (1, -1)$ with a = 1, 2,

3. Each left-handed neutrino $N_{aL}=(N_{aR})^c$ implies a new right-handed neutrino beyond the SM. The three Higgs triplets are $\rho = (\rho_1^+, \rho^0, \rho_2^+)^T \sim (3, \frac{2}{3}), \ \eta = (\eta_1^0, \eta^-, \eta_2^0)^T \sim (3, -\frac{1}{3}),$ and $\chi = (\chi_1^0, \chi^-, \chi_2^0)^T \sim (3, -\frac{1}{3})$. The necessary vacuum expectation values for generating all tree-level quark masses are $\langle \rho \rangle = (0, \frac{v_1}{\sqrt{2}}, 0)^T, \ \langle \eta \rangle = (\frac{v_2}{\sqrt{2}}, 0, 0)^T, \ \text{and} \ \langle \chi \rangle = (0, 0, \frac{w}{\sqrt{2}})^T.$ Gauge bosons in this model get masses through the covariant kinetic term of the Higgs bosons,

$$\mathcal{L}^{H} = \sum_{H=\chi,\eta,
ho} (D_{\mu}H)^{\dagger}(D_{\mu}H),$$

where the covariant derivative for the electroweak symmetry is defined as

$$D_{\mu} = \partial_{\mu} - igW_{\mu}^{a}T^{a} - g_{X}T^{9}XX_{\mu}, \quad a = 1, 2, ..., 8,$$
 (1)

and $T^9 \equiv \frac{I_3}{\sqrt{6}}$ and $\frac{1}{\sqrt{6}}$ for (anti)triplets and singlets [18]. It can be identified that

$$g = es_W, \qquad \frac{g_X}{g} = \frac{3\sqrt{2}s_W}{\sqrt{3 - 4s_W^2}},$$
 (2)

where e and s_W are, respectively, the electric charge and sine of the Weinberg angle, $s_W^2 \simeq 0.231$.

The model includes two pairs of singly charged gauge bosons, denoted as W^{\pm} and Y^{\pm} , defined as

$$W_{\mu}^{\pm} = \frac{W_{\mu}^{1} \mp iW_{\mu}^{2}}{\sqrt{2}}, \qquad m_{W}^{2} = \frac{g^{2}}{4}(v_{1}^{2} + v_{2}^{2}),$$

$$Y_{\mu}^{\pm} = \frac{W_{\mu}^{6} \pm iW_{\mu}^{7}}{\sqrt{2}}, \qquad m_{Y}^{2} = \frac{g^{2}}{4}(w^{2} + v_{1}^{2}). \tag{3}$$

The bosons W^{\pm} are identified with the SM ones, leading to $v_1^2 + v_2^2 \equiv v^2 = (246 \text{ GeV})^2$. In the remainder of the text, we will consider in detail the simple case $v_1 = v_2 = v/\sqrt{2} = \sqrt{2}m_W/g$ given in Refs. [6,19].

The two global symmetries—namely, the normal and new lepton numbers denoted, respectively, as L and \mathcal{L} —were introduced. They are related to each other by [16,20] $L = \frac{4}{\sqrt{3}}T_8 + \mathcal{L}$. The detailed values of nonzero lepton numbers L and \mathcal{L} are listed in Table I.

TABLE I. Nonzero lepton number L (left) and \mathcal{L} (right) of leptons and Higgs bosons in the 331RHN.

Fields	N_L	$ u_L$	e_L	e_R	ρ_2^+	η_2^0	χ_1^0	χ^{-}
\overline{L}	-1	1	1	1	-2	-2	2	2
Fields	χ		η	ρ		ψ_{aL}		e_{aR}
\mathcal{L}	$\frac{4}{3}$		<u>2</u> 3	$\frac{2}{3}$		<u>1</u> 3		1

All tree-level lepton mass terms come from the following Yukawa part:

$$\mathcal{L}_{l}^{Y} = -h_{ab}^{e} \overline{\psi_{aL}} \rho e_{bR} + h_{ab}^{\nu} \epsilon^{ijk} \overline{(\psi_{aL})_{i}} (\psi_{bL})_{i}^{c} \rho_{k}^{*} + \text{H.c.}, \quad (4)$$

where e^{ijk} is the antisymmetric tensor $e^{123} = 1$, $(\psi_{aL})^c \equiv ((\nu_{aL})^c, (e_{aL})^c, (N_{aL})^c)^T$, and h^{ν} is an antisymmetric matrix, $h^{\nu}_{ab} = -h^{\nu}_{ba}$. The first term of Eq. (4) generates charged lepton masses m_a satisfying $h^e_{ab} \equiv \sqrt{2}\delta_{ab}m_a/v_1$ in order to avoid LFV processes at the tree level. The second term in Eq. (4) is expanded as follows:

$$h_{ab}^{\nu} \epsilon^{ijk} \overline{(\psi_{aL})_i} (\psi_{bL})_j^c \rho_k^* = 2h_{ab}^{\nu} \left[-\overline{e_{aL}} (\nu_{bL})^c \rho_2^- - \overline{\nu_{aL}} (N_{bL})^c \rho^{0*} + \overline{e_{aL}} (\nu_{bL})^c \rho_1^- \right],$$

$$(5)$$

where we have used the equality $\overline{N_{aL}}(\nu_{bL})^c = \overline{\nu_{bL}}(N_{aL})^c$,... The second term on the left-hand side of Eq. (5) contributes a Dirac neutrino mass term $-\mathcal{L}_{\text{mass}}^{\nu} = \overline{\nu_L} m_D N_R + \text{H.c.}$, where $\nu_L \equiv (\nu_{1L}, \nu_{2L}, \nu_{3L})^T$, $N_R \equiv ((N_{1L})^c, (N_{3L})^c)^T$, and $(m_D)_{ab} \equiv \sqrt{2} v_1 h_{ab}^{\nu}$. The model can predict a neutrino mass spectrum that is consistent with current neutrino data [21] when loop corrections are included, where all new neutrinos are very light [16]. As a result, they will give suppressed LFV decay rates.

Now we consider a 331ISS model as an extension of the above 331RHN model, where three right-handed neutrinos [which are gauge singlets, $X_{aR} \sim (1,0)$, a=1,2,3] are added. Now tree-level neutrino masses and mixing angles arise from the ISS mechanism. Requiring that \mathcal{L} is only softly broken, the additional Yukawa part is

$$-\mathcal{L}_{X_R} = Y_{ab}\overline{\psi_{aL}}\chi X_{bR} + \frac{1}{2}(\mu_X)_{ab}\overline{(X_{aR})^c}X_{bR} + \text{H.c.}, \quad (6)$$

where μ_X is a 3×3 symmetric matrix and $L(X_{aR}) = \mathcal{L}(X_{aR}) = -1$. The last term in Eq. (6) is the only one that violates both L and \mathcal{L} , and hence it can be assumed to be small, which is exactly the case in the ISS models. The first term generates mass for heavy neutrinos, resulting in a large Yukawa coupling Y_{ab} with $SU(3)_L$ Higgs triplets. In addition, the ISS mechanism allows for large entries in the Dirac mass matrix m_D originated from Eq. (4), which is the opposite of the well-known requirement in the 331RHN model.

In the basis $\nu_L' = (\nu_L, N_L, (X_R)^c)^T$ and $(\nu_L')^c = ((\nu_L)^c, (N_L)^c, X_R)^T$, Eqs. (4) and (6) give a neutrino mass term corresponding to a block form of the mass matrix, namely,

$$-L_{\text{mass}}^{\nu} = \frac{1}{2} \overline{\nu_L} M^{\nu} (\nu_L^{\prime})^c + \text{H.c.}, \quad \text{where}$$

$$M^{\nu} = \begin{pmatrix} 0 & m_D & 0 \\ m_D^T & 0 & M_R \\ 0 & M_R^T & \mu_X \end{pmatrix}, \tag{7}$$

where M_R is a 3×3 matrix $(M_R)_{ab} \equiv Y_{ab} \frac{w}{\sqrt{2}}$ with a, b=1, 2, 3. Neutrino sub-bases are denoted as $\nu_R = ((\nu_{1L})^c, (\nu_{2L})^c, (\nu_{3L})^c)^T$, $N_R = ((N_{1L})^c, (N_{2L})^c, (N_{3L})^c)^T$, and $X_L = ((X_{1R})^c, (X_{2R})^c, (X_{3R})^c)^T$.

The matrix M^{ν} can be written in the normal seesaw form,

$$\begin{split} M^{\nu} &= \begin{pmatrix} 0 & M_D \\ M_D^T & M_N \end{pmatrix}, \quad \text{where } M_D \equiv (m_D, 0), \quad \text{and} \\ M_N &= \begin{pmatrix} 0 & M_R \\ M_D^T & \mu_X \end{pmatrix}. \end{split} \tag{8}$$

The mass matrix M^{ν} is diagonalized by a 9×9 unitary matrix U^{ν} [17,22],

$$U^{\nu T} M^{\nu} U^{\nu} = \hat{M}^{\nu} = \operatorname{diag}(m_{n_1}, m_{n_2}, ..., m_{n_9}) = \operatorname{diag}(\hat{m}_{\nu}, \hat{M}_N),$$
(9)

where m_{n_i} (i=1,2,...,9) are mass eigenvalues of the nine mass eigenstates n_{iL} (i.e., physical states of neutrinos), $\hat{m}_{\nu} = \mathrm{diag}(m_{n_1}, m_{n_2}, m_{n_3})$, and $\hat{M}_N = \mathrm{diag}(m_{n_4}, m_{n_5}, ..., m_{n_9})$. They correspond to the masses of the three active neutrinos n_{aL} (a=1,2,3) and six extra neutrinos n_{IL} (I=4,5,...,9). The relations between the flavor and mass eigenstates are

$$\nu'_L = U^{\nu*} n_L$$
, and $(\nu'_L)^c = U^{\nu} (n_L)^c$, (10)

where $n_L \equiv (n_{1L}, n_{2L}, ..., n_{9L})^T$ and $(n_L)^c \equiv ((n_{1L})^c, (n_{2L})^c, ..., (n_{9L})^c)^T$.

A four-component (Dirac) spinor n_i is defined as $n_i \equiv (n_{iL}, (n_{iL})^c)^T = n_i^c = (n_i)^c$, where the chiral components are $n_{L,i} \equiv P_L n_i$ and $n_{R,i} \equiv P_R n_i = (n_{L,i})^c$ with chiral operators $P_{L,R} = \frac{1 \pm \gamma_5}{2}$. Similarly, the definitions for the original neutrino states are $\nu_a \equiv (\nu_{L,a}, (\nu_{L,a})^c)^T$, $\nu_a \equiv (N_{L,a}, (N_{L,a})^c)^T$, $X_I \equiv ((X_{R,I})^c, X_{R,I})^T$, and $\nu' = (\nu, N)^T$. The relations in Eq. (10) can be written as follows:

$$P_L \nu'_i = \nu'_{i,L} = U^{\nu*}_{ij} n_{jL},$$
 and $P_R \nu'_i = \nu'_{iR} = U^{\nu}_{ij} n_{jR},$ $i, j = 1, 2, ..., 9.$ (11)

In general, U^{ν} is written in the form [23]

$$U^{\nu} = \Omega \begin{pmatrix} U & \mathbf{O} \\ \mathbf{O} & V \end{pmatrix}, \tag{12}$$

where **O** is a 3×6 null matrix, and U, V, and Ω are 3×3 , 6×6 , and 9×9 unitary matrices, respectively. Ω can be formally written as

$$\Omega = \exp\begin{pmatrix} \mathbf{O} & R \\ -R^{\dagger} & \mathbf{O} \end{pmatrix} = \begin{pmatrix} 1 - \frac{1}{2}RR^{\dagger} & R \\ -R^{\dagger} & 1 - \frac{1}{2}R^{\dagger}R \end{pmatrix} + \mathcal{O}(R^{3}),$$
(13)

where R is a 3 × 6 matrix with the maximal absolute values for all entries |R| satisfying |R| < 1. The matrix $U = U_{\text{PMNS}}$ is the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix [24],

$$U_{\text{PMNS}} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix} \operatorname{diag}(1, e^{i\frac{\sigma}{2}}, e^{i\frac{\beta}{2}}), \tag{14}$$

and $c_{ab} \equiv \cos \theta_{ab}$, $s_{ab} \equiv \sin \theta_{ab}$. The Dirac phase δ and Majorana phases α , β are fixed as $\delta = \pi$, $\alpha = \beta = 0$. In the normal hierarchy scheme, the best-fit values of neutrino oscillation parameters are given as [21]

$$\Delta m_{21}^2 = 7.370 \times 10^{-5} \text{ eV}^2,$$

$$\Delta m^2 = 2.50 \times 10^{-3} \text{ eV}^2, \qquad s_{12}^2 = 0.297,$$

$$s_{23}^2 = 0.437, \qquad s_{13}^2 = 0.0214, \qquad (15)$$

where $\Delta m_{21}^2 = m_{n_2}^2 - m_{n_1}^2$ and $\Delta m^2 = m_{n_3}^2 - \frac{\Delta m_{21}^2}{2}$. The condition $v_1 \ll w$ gives the reasonable condition $|M_D| \ll |M_N|$, where $|M_D|$ and $|M_N|$ denote the characteristic scales of M_D and M_N . Hence, the following seesaw relations are valid [23]:

$$R^* \simeq (-m_D M^{-1}, m_D (M_R^T)^{-1}),$$
 (16)

$$m_D M^{-1} m_D^T \simeq m_\nu \equiv U_{\text{PMNS}}^* \hat{m}_\nu U_{\text{PMNS}}^\dagger,$$
 (17)

$$V^* \hat{M}_N V^{\dagger} \simeq M_N + \frac{1}{2} R^T R^* M_N + \frac{1}{2} M_N R^{\dagger} R,$$
 (18)

where

$$M \equiv M_R \mu_X^{-1} M_R^T. \tag{19}$$

In the model under consideration, the Dirac neutrino mass matrix m_D must be antisymmetric. Equivalently, m_D has only three independent parameters x_{12} , x_{13} , and z,

$$m_D \equiv z \begin{pmatrix} 0 & x_{12} & x_{13} \\ -x_{12} & 0 & 1 \\ -x_{13} & -1 & 0 \end{pmatrix}, \tag{20}$$

where $z = \sqrt{2}v_1 h_{23}^{\nu}$. In contrast, the matrix m_{ν} in Eq. (17) is symmetric, $(m_{\nu})_{ij} = (m_{\nu})_{ji}$, implying that

$$0 = (m_{\nu})_{ij} - (m_{\nu})_{ji} \sim x_{12}[(M^{-1})_{12} - (M^{-1})_{21}]$$

+ $x_{13}[(M^{-1})_{13} - (M^{-1})_{31}] + (M^{-1})_{23} - (M^{-1})_{32},$

with i, j = 1, 2, 3. This means that a symmetric matrix M will give a right antisymmetric matrix m_D . To fit the neutrino data, there must exist matrices M and m_D that satisfy the first equality in Eq. (17). Here we choose M to be symmetric for simplicity. There must exist some sets of z, x_{12}, x_{13} , and M_{ij} ($i \le j \le 3$) that satisfy the six equations $(m_D M^{-1} m_D^T)_{ij} = (m_\nu)_{ij}$, with $i \le j \le 3$. From the three equations corresponding to i = j = 1, 2, 3, we can write $(M^{-1})_{ii}$ as three functions of z, x_{12}, x_{13} , and $(M^{-1})_{ij}$ ($i \ne j$). Inserting them into the three remaining equalities, and taking some intermediate steps, we obtain

$$-(m_{\nu})_{13}x_{12} + (m_{\nu})_{12}x_{13} = (m_{\nu})_{11},$$

$$-(m_{\nu})_{23}x_{12} + (m_{\nu})_{22}x_{13} = (m_{\nu})_{12},$$

$$-(m_{\nu})_{33}x_{12} + (m_{\nu})_{23}x_{13} = (m_{\nu})_{13},$$
(21)

where we exclude the case of x_{12} , $x_{13} = 0$. Solving the above three equations leads to two solutions for $x_{12,13}$ and a strict relation among $(m_{\nu})_{ij}$:

$$x_{12} = \frac{(m_{\nu})_{11}(m_{\nu})_{23} - (m_{\nu})_{13}(m_{\nu})_{12}}{(m_{\nu})_{12}(m_{\nu})_{33} - (m_{\nu})_{13}(m_{\nu})_{23}},$$

$$x_{13} = \frac{(m_{\nu})_{11}(m_{\nu})_{33} - (m_{\nu})_{13}^{2}}{(m_{\nu})_{12}(m_{\nu})_{33} - (m_{\nu})_{13}(m_{\nu})_{23}},$$

$$0 = (m_{\nu})_{11}(m_{\nu})_{23}^{2} + (m_{\nu})_{22}(m_{\nu})_{13}^{2} + (m_{\nu})_{33}(m_{\nu})_{12}^{2}$$

$$- (m_{\nu})_{11}(m_{\nu})_{22}(m_{\nu})_{33} - 2(m_{\nu})_{12}(m_{\nu})_{13}(m_{\nu})_{23}.$$
(22)

Interestingly, the last relation in Eq. (22) allows us to predict possible values of the unknown neutrino mass based on the identification given in Eq. (17). Using the experimental data given in Eq. (15), we derive that $m_{\nu_1}=0$ in the normal hierarchy scheme. The Dirac matrix now only depends on z:

$$m_D \simeq z \times \begin{pmatrix} 0 & 0.545 & 0.395 \\ -0.545 & 0 & 1 \\ -0.395 & -1 & 0 \end{pmatrix}$$
. (23)

The above discussion also gives $M = \text{diag}(10^{10}z^2, 7.029 \times 10^{10}z^2, -2.377 \times 10^{11}z^2)$ for a diagonal M_R . In

this work, we also consider the simple case where M_R is diagonal and all elements are positive. We also assume that $|m_{\nu}| < \mu_X \ll |m_D| < |M_R|$. We then derive that heavy neutrino masses are approximately equal to the entries of M_R , as given in Eq. (19). However, this approximation is not good for investigating LFVHDs, where a divergent cancellation in the numerical computation is strictly required. Instead, we will use the numerical solutions of heavy neutrino masses as well as the mixing matrix U^{ν} so that a total divergent part vanishes in the final numerical results. This treatment will avoid unphysical contributions originated from divergent parts.

Another parametrization shown in Ref. [8] can be applied to the general cases of nonzero δ as well as both the inverse and normal hierarchy cases of active neutrino masses. With the aim of finding regions with large LFVHDs, we will choose the simple case of m_D given in Eq. (23).

For simplicity in the numerical study, we will consider the diagonal matrix M_R in the degenerate case $M_R = M_{R_1} = M_{R_2} = M_{R_3} \equiv k \times z$. The parameter k will be fixed at small values that result in large LFVHD effects. The total neutrino mass matrix in Eq. (7) depends on only the free parameter z. The heavy neutrino masses and the matrix U^{ν} can be solved numerically, which is not affected by z because $|\mu_X| \ll z$.

Using the exact numerical solutions for the neutrino masses and mixing matrix U^{ν} for our investigation, we emphasize that the masses and mixing parameters of active neutrinos derived from the numerical diagonalization of the matrix M^{ν} given in Eq. (7) should satisfy the 3σ constraint of the experimental data. In contrast, neutrino masses and mixing parameters defining the matrix m_{ν} in Eq. (17), which are used to calculate th matrix m_D , are considered as free parameters. In other words, the experimental values of neutrino masses and mixing parameters are only used to estimate the allowed ranges of free parameters determining the mass matrix M^{ν} . After that, it is diagonalized numerically to find the neutrino masses as well as the mixing matrix U^{ν} . The mixing parameters will be calculated from the matrix $U_{\rm PMNS}$, which is related to U^{ν} by the relation (12). Requiring that the expansion of Ω in Eq. (13) and the ISS condition $|\mu_X| > m_{n_3}$ are valid, we find that small values of k > 1 are allowed. In particular, we find that if three mixing parameters are fixed at the three respective center values, the two inputs for the active neutrino masses may be outside of (but very close to) the 3σ ranges with k = 5. When $k \ge 5.5$, we always find that the input lies within the 3σ ranges of the experimental data that produces the consistent numerical solutions of active neutrino masses. When $k \ge 9$, the input corresponding to all center values given in Eq. (15) always produces numerical solutions lying in the 3σ ranges of experimental data.

The LFVHD rates depend strongly on the unitarity of the mixing matrix U^{ν} and heavy neutrino masses. On the other hand, they are weakly affected by the requirement that solutions for active neutrino masses and mixing parameters

satisfy the 3σ experimental data. Hence, we will use the matrix m_D given in Eq. (23) and $k \ge 5.5$ for our numerical investigation. We numerically checked that our choice produces reasonable values for the neutrino data close to the 3σ ranges mentioned above.

B. Higgs and gauge bosons

To study the LFVHD effects, we will choose the simple case of the Higgs potential discussed in Refs. [6,19], namely,

$$\mathcal{V} = \mu_1^2 (\rho^{\dagger} \rho + \eta^{\dagger} \eta) + \mu_2^2 \chi^{\dagger} \chi + \lambda_1 (\rho^{\dagger} \rho + \eta^{\dagger} \eta)^2 + \lambda_2 (\chi^{\dagger} \chi)^2$$
$$+ \lambda_{12} (\rho^{\dagger} \rho + \eta^{\dagger} \eta) (\chi^{\dagger} \chi) - \sqrt{2} f(\varepsilon_{ijk} \eta^i \rho^j \chi^k + \text{H.c.}), \quad (24)$$

where f is a mass parameter and is assumed to be real. The detailed calculations for finding the masses and mass eigenstates of Higgs bosons were presented in Refs. [6,19], where the minimum condition results in $v_1 = v_2$. Here we will only list the part that is involved in LFVHDs.

The model contains two pairs of singly charged Higgs bosons $H_{1,2}^\pm$ and Goldstone bosons of the gauge bosons W^\pm and Y^\pm , which are denoted as G_W^\pm and G_Y^\pm , respectively. The masses of all charged Higgs bosons are $m_{H_2^\pm}^2 = fw(t_\theta^2+1)$, $m_{H_2^\pm}^2 = 2fw$, and $m_{G_W^\pm}^2 = m_{G_Y^\pm}^2 = 0$, where $t_\theta = v_2/w$. The relations between the original and mass eigenstates of the charged Higgs bosons are

$$\begin{pmatrix} \rho_{1}^{\pm} \\ \eta^{\pm} \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} -1 & 1 \\ 1 & 1 \end{pmatrix} \begin{pmatrix} G_{W}^{\pm} \\ H_{2}^{\pm} \end{pmatrix},$$

$$\begin{pmatrix} \rho_{2}^{\pm} \\ \gamma^{\pm} \end{pmatrix} = \begin{pmatrix} -s_{\theta} & c_{\theta} \\ c_{\theta} & s_{\theta} \end{pmatrix} \begin{pmatrix} G_{Y}^{\pm} \\ H_{1}^{\pm} \end{pmatrix}.$$
(25)

The neutral scalars are expanded as

$$\rho^{0} = \frac{1}{\sqrt{2}} (v_{1} + S_{1} + iA_{1}), \qquad \eta_{1}^{0} = \frac{1}{\sqrt{2}} (v_{2} + S_{2} + iA_{2}),$$

$$\chi_{2}^{0} = \frac{1}{\sqrt{2}} (w + S_{3}' + iA_{3}'),$$

$$\eta_{2}^{0} = \frac{1}{\sqrt{2}} (S_{2}' + iA_{2}'), \qquad \chi_{1}^{0} = \frac{1}{\sqrt{2}} (S_{3} + iA_{3}). \tag{26}$$

There are four physical CP-even Higgs bosons $h_{1,2,3,4}^0$ and a Goldstone boson of the non-Hermitian gauge boson. The neutral Higgs components relevant for this work are defined via

$$\begin{pmatrix} S_1 \\ S_2 \\ S_3' \end{pmatrix} = \begin{pmatrix} -\frac{c_{\alpha}}{\sqrt{2}} & \frac{s_{\alpha}}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ -\frac{c_{\alpha}}{\sqrt{2}} & \frac{s_{\alpha}}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \\ s_{\alpha} & c_{\alpha} & 0 \end{pmatrix} \begin{pmatrix} h_1^0 \\ h_2^0 \\ h_3^0 \end{pmatrix}, \quad (27)$$

where $s_{\alpha} = \sin \alpha$ and $c_{\alpha} = \cos \alpha$, and they are defined by

TABLE II. Couplings related to the SM-like Higgs decay $h_1^0 \rightarrow e_a e_b$ in the 331ISS model. All momenta in the Feynman rules corresponding to these vertices are incoming.

Vertex	Coupling		
$h_1^0 \overline{e_a} e_a$	$rac{igm_a}{2m_W}C_{lpha}$		
$h_1^0 \overline{n_i} n_j$	$rac{igc_{a}}{2m_{W}}(\lambda_{ij}^{0}P_{L}+\lambda_{ij}^{0st}P_{R})$		
$H_1^+\overline{n_i}e_b,H_1^-\overline{e_a}n_i$	$-rac{igc_{ heta}}{m_{W}}(\lambda_{bi}^{L,1}P_{L}+\lambda_{bi}^{R,1}P_{R}),-rac{igc_{ heta}}{m_{W}}(\lambda_{ai}^{L,1*}P_{R}+\lambda_{ai}^{R,1*}P_{L})$		
$H_2^+\overline{n_i}e_b,H_2^-\overline{e_a}n_i$	$-rac{ig}{\sqrt{2}m_{w}}(\lambda_{bi}^{L,2}P_{L}+\lambda_{bi}^{R,2}P_{R}),-rac{ig}{\sqrt{2}m_{w}}(\lambda_{ai}^{L,2*}P_{R}+\lambda_{ai}^{R,2*}P_{L})$		
$W^+_\mu \overline{n_i} e_b, \ W^\mu \overline{e_a} n_i$	$rac{ig}{\sqrt{2}}U_{bi}^{ u}\gamma^{\mu}P_{L}, rac{ig}{\sqrt{2}}U_{ai}^{ ust}\gamma^{\mu}P_{L}$		
$Y_{\mu}^{+}\overline{n_{i}}e_{b},\ Y_{\mu}^{-}\overline{e_{a}}n_{i}$	$rac{ig}{\sqrt{2}}U^ u_{(b+3)i}\gamma^\mu P_L, rac{ig}{\sqrt{2}}U^{ ust}_{(a+3)i}\gamma^\mu P_L$		
$H_1^+ h_1^0 Y_\mu^-, \ Y_\mu^+ H_1^- h_1^0$	$rac{ig}{2\sqrt{2}}(c_{lpha}c_{ heta}+\sqrt{2}s_{lpha}s_{ heta})(p_{h_{1}^{0}}-p_{H_{1}^{+}})^{\mu},rac{ig}{2\sqrt{2}}(c_{lpha}c_{ heta}+\sqrt{2}s_{lpha}s_{ heta})(p_{H_{1}^{-}}-p_{h_{1}^{0}})^{\mu}$		
$h_1^0 W_\mu^+ W_ u^-$	$-igm_Wc_{lpha}g^{\mu u}$		
$h_1^0 Y_\mu^+ Y_ u^-$	$\frac{igm_Y}{\sqrt{2}}(\sqrt{2}s_{lpha}c_{ heta}-c_{lpha}s_{ heta})g^{\mu u}$		
$h_1^0 H_1^+ H_1^-$	$i\lambda_{H_1}^{\pm} = -iw[s_{\alpha}c_{\theta}^2\lambda_{12} + 2s_{\alpha}s_{\theta}^2\lambda_2 - \sqrt{2}(2c_{\alpha}c_{\theta}^2\lambda_1 + c_{\alpha}s_{\theta}^2\lambda_{12})t_{\theta} - \frac{\sqrt{2}}{v_3}fc_{\alpha}c_{\theta}s_{\theta}]$		
$h_1^0 H_2^+ H_2^-$	$i\lambda_{H_2}^{\pm} = -iv_1(-2\sqrt{2}c_a\lambda_1 + \frac{s_av_3\lambda_{12} + s_af}{v_1})$		

$$s_{\alpha} = \frac{(4\lambda_{1} - m_{h_{1}^{0}}^{2}/v_{2}^{2})t_{\theta}}{r}, \qquad c_{\alpha} = \frac{\sqrt{2}(\lambda_{12} - \frac{f}{w})}{r},$$

$$r = \sqrt{2\left(\lambda_{12} - \frac{f}{w}\right)^{2} + (4\lambda_{1} - m_{h_{1}^{0}}^{2}/v_{2}^{2})^{2}t_{\theta}^{2}}.$$
(28)

There is one neutral CP-even Higgs boson h_1^0 with a mass proportional to the electroweak scale,

$$m_{h_1^0}^2 = \frac{w^2}{2} \left[4\lambda_1 t_{\theta}^2 + 2\lambda_2 + \frac{f t_{\theta}^2}{w} - \sqrt{\left(2\lambda_2 + \frac{f t_{\theta}^2}{w} - 4\lambda_1 t_{\theta}^2\right)^2 + 8t_{\theta}^2 \left(\frac{f}{w} - \lambda_{12}\right)^2} \right]. \tag{29}$$

The decoupling limit $t_{\theta} \ll 1$ $(v_1 \ll w)$ gives $m_{h_1^0}^2 \sim \mathcal{O}(m_W^2)$ and $s_{\alpha} \simeq 0$ [19], resulting in the couplings similar to those predicted by the SM; see Table II. Hence, h_1^0 is identified with the SM-like Higgs boson found at the LHC.

III. COUPLINGS AND ANALYTIC FORMULAS INVOLVED WITH LFVHDS

A. Couplings

In this section we present Yukawa couplings in terms of U^{ν} and physical neutrino masses. From this, amplitudes and the LFVHD rate are formulated in terms of physical

masses and mixing parameters. The equality derived from Eq. (9), $M^{\nu} = U^{\nu*} \hat{M}^{\nu} U^{\nu\dagger}$, gives

$$\begin{split} M_{ab}^{\nu} &= (U^{\nu*}\hat{M}^{\nu}U^{\nu\dagger})_{ab} = 0 \rightarrow U_{ak}^{\nu*}U_{bk}^{\nu*}m_{n_{k}} = 0, \\ \sqrt{2}v_{1}h_{ab}^{\nu} &= (m_{D})_{ab} = (M^{\nu})_{a(b+3)} \\ &= (U^{\nu*}\hat{M}^{\nu}U^{\nu\dagger})_{a(b+3)} = U_{ak}^{\nu*}U_{(b+3)k}^{\nu*}m_{n_{k}}, \\ \frac{w}{\sqrt{2}}Y_{ab} &= (M_{R})_{ab} = (M^{\nu})_{(a+3)(b+6)} \\ &= U_{(a+3)k}^{\nu*}U_{(b+6)k}^{\nu*}m_{n_{k}}, \end{split} \tag{30}$$

where a, b=1, 2, 3, and the sum is taken over k=1,2,...,9.

The relevant couplings in the first term of the Lagrangian (4) are

$$\begin{split} &-h_{ab}^{e}\overline{\psi_{aL}}\rho e_{bR} + \text{H.c.} \\ &= -\frac{gm_{a}}{m_{W}}[\overline{\nu_{aL}}e_{aR}\rho_{1}^{+} + \overline{e_{aL}}e_{aR}\rho^{0} + \overline{N_{aL}}e_{aR}\rho_{2}^{+} + \text{H.c.}] \\ &\supset \frac{gm_{a}c_{a}}{2m_{W}}h_{1}^{0}\overline{e_{a}}e_{a} \\ &- \frac{gm_{a}}{m_{W}}[c_{\theta}(U_{(a+3)i}^{\nu}\overline{n_{i}}P_{R}e_{a}H_{1}^{+} + U_{(a+3)i}^{\nu*}\overline{e_{a}}P_{L}n_{i}H_{1}^{-})] \\ &- \frac{gm_{a}}{\sqrt{2}m_{W}}[(U_{ai}^{\nu}\overline{n_{i}}P_{R}e_{a}H_{2}^{+} + U_{ai}^{\nu*}\overline{e_{a}}P_{L}n_{i}H_{2}^{-})]. \end{split}$$
(31)

The relevant couplings in the second term of the Lagrangian (4) are

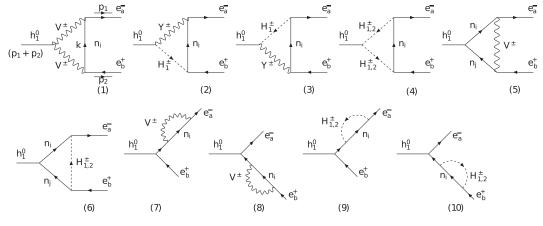


FIG. 1. One-loop Feynman diagrams contributing to the decay $h_1^0 \to e_a e_b$ in the unitary gauge. Here $V^{\pm} = W^{\pm}, Y^{\pm}$.

$$\begin{split} h_{ab}^{\nu} \epsilon^{ijk} \overline{(\psi_{aL})_{i}} (\psi_{bL})_{j}^{c} \rho_{k}^{*} + \text{H.c.} \\ &= 2h_{ab}^{\nu} \left[-\overline{e_{aL}} (\nu_{bL})^{c} \rho_{2}^{-} - \overline{\nu_{aL}} (N_{bL})^{c} \rho^{0*} + \overline{e_{aL}} (\nu_{bL})^{c} \rho_{1}^{-} \right] \\ &= \frac{gc_{\alpha}}{2m_{W}} h_{1}^{0} \left[\sum_{c=1}^{3} U_{ci}^{\nu} U_{cj}^{\nu*} \overline{n_{i}} (m_{n_{i}} P_{L} + m_{n_{j}} P_{R}) n_{j} \right] \\ &- \frac{gc_{\theta}}{m_{W}} \left[(m_{D})_{ab} U_{bi}^{\nu} H_{1}^{-} \overline{e_{a}} P_{R} n_{i} + \text{H.c.} \right] \\ &+ \frac{g}{\sqrt{2}m_{W}} \left[(m_{D})_{ab} U_{(b+3)i}^{\nu} H_{2}^{-} \overline{e_{a}} P_{R} n_{i} + \text{H.c.} \right], \end{split}$$
(32)

where the last line is derived following the calculation in Ref. [22]: $\overline{\nu_L} M_D((N_L)^c, X_R)^T \leftrightarrow \overline{\nu_{aL}} (M_D)_{aI} N_{IR}$. The first term in Eq. (6) gives the following couplings:

$$-Y_{ab}\overline{\psi_{aL}}\chi X_{bR} + \text{H.c.}$$

$$= -\frac{\sqrt{2}}{w}(M_R)_{ab}[\overline{\nu_{aL}}\chi_1^0 + \overline{e_{aL}}\chi^- + \overline{N_{aL}}\chi_2^0]X_{bR} + \text{H.c.}$$

$$\supset -\frac{gt_{\theta}}{\sqrt{2}m_W}(M_R)_{ab}[s_{\alpha}U^{\nu}_{(a+3)i}U^{\nu}_{(b+6)j}\overline{n_i}P_Rn_jh_1^0$$

$$+\sqrt{2}s_{\theta}U^{\nu}_{(b+6)i}\overline{e_a}P_Rn_iH_1^- + \text{H.c.}], \tag{33}$$

where we have used $t_{\theta} = v_1/w \rightarrow 1/w = t_{\theta}/v_1 = gt_{\theta}/(\sqrt{2}m_W)$. The LFVHD couplings between leptons and charged gauge bosons (W^{\pm}, Y^{\pm}) are

$$\mathcal{L}^{\ell\ell V} = \overline{\psi_{aL}} \gamma^{\mu} D_{\mu} \psi_{aL}$$

$$\supset \frac{g}{\sqrt{2}} (\overline{e_{aL}} \gamma^{\mu} \nu_{aL} W_{\mu}^{-} + \overline{e_{aL}} \gamma^{\mu} N_{aL} Y_{\mu}^{-}) + \text{H.c.}$$

$$= \frac{g}{\sqrt{2}} [U_{ai}^{\nu*} \overline{e_{a}} \gamma^{\mu} P_{L} n_{i} W_{\mu}^{-} + U_{ai}^{\nu} \overline{n_{i}} \gamma^{\mu} P_{L} e_{a} W_{\mu}^{+}$$

$$+ U_{(a+3)i}^{\nu*} \overline{e_{a}} \gamma^{\mu} P_{L} n_{i} Y_{\mu}^{-} + U_{(a+3)i}^{\nu} \overline{n_{i}} \gamma^{\mu} P_{L} e_{a} Y_{\mu}^{+}], \quad (34)$$

where $D_{\mu} = \partial_{\mu} - \frac{ig}{2}(W_{\mu}^{a}\lambda^{a} + t \times (-\frac{1}{3})B_{\mu})$, λ_{a} (a = 1, 2, ..., 8) are the Gell-Mann matrices, and $t = g_{X}/g$. The charged gauge bosons are $W_{\mu}^{\pm} = \frac{W_{\mu}^{1} \mp iW_{\mu}^{2}}{\sqrt{2}}$ and $Y_{\mu}^{\pm} = \frac{W_{\mu}^{6} \pm iW_{\mu}^{7}}{\sqrt{2}}$.

By defining a symmetric coefficient $\lambda_{ij}^0 = \lambda_{ji}^0$, namely,

$$\begin{split} \lambda_{ij}^0 &= \sum_{c=1}^3 \left(U_{ci}^\nu U_{cj}^{\nu*} m_{n_i} + U_{ci}^{\nu*} U_{cj}^\nu m_{n_j} \right) \\ &- \sum_{c,d=1}^3 \sqrt{2} \, t_\alpha t_\theta (M_R^*)_{cd} \bigg[U_{(c+3)i}^{\nu*} U_{(d+6)j}^{\nu*} + U_{(c+3)j}^{\nu*} U_{(d+6)i}^{\nu*} \bigg], \end{split}$$

the coupling $h_1^0 \bar{n}_i n_j$ derived from Eqs. (32) and (33) is written in the symmetric form $\frac{gc_\alpha}{4m_W}h_1^0 \overline{n}_i [\lambda_{ij}^0 P_L + \lambda_{ij}^{0*} P_R] n_j$, which gives the right vertex coupling based on the Feynman rules given in Ref. [25]. The Yukawa couplings of charged Higgs bosons are defined by

$$\lambda_{ai}^{R,1} = m_a U_{(a+3)i}^{\nu},$$

$$\lambda_{ai}^{L,1} = \sum_{c=1}^{3} \left[(m_D^*)_{ac} U_{ci}^{\nu*} + t_{\theta}^2 (M_R^*)_{ac} U_{(c+6)i}^{\nu*} \right],$$

$$\lambda_{ai}^{R,2} = m_a U_{ai}^{\nu}, \qquad \lambda_{ai}^{L,2} = -\sum_{c=1}^{3} (m_D^*)_{ac} U_{(c+3)i}^{\nu*}. \tag{35}$$

Finally, all of the couplings involved in LFV processes are listed in Table II. The model predicts that the following couplings are zero: $h_1^0W^{\pm}Y^{\mp}$, $h_1^0W^{\pm}H_{1,2}^{\mp}$, $h_1^0Y^{\pm}H_2^{\mp}$, and $h_1^0H_1^{\pm}H_2^{\mp}$.

B. Analytic formulas

The effective Lagrangian of the LFVHDs of the SM-like Higgs boson $h_1^0 \to e_a^\pm e_b^\mp$ is

$$\mathcal{L}^{\mathrm{LFVH}} = h_1^0(\Delta_{(ab)L}\overline{e_a}P_Le_b + \Delta_{(ab)R}\overline{e_a}P_Re_b) + \mathrm{H.c.},$$

where the scalar factors $\Delta_{(ab)L,R}$ arise from the loop contributions. In the unitary gauge, the one-loop Feynman diagrams contributing to this LFVHD amplitude are shown in Fig. 1.

The partial width of the decay is

$$\Gamma(h_1^0 \to e_a e_b) \equiv \Gamma(h_1^0 \to e_a^- e_b^+) + \Gamma(h_1^0 \to e_a^+ e_b^-)$$

$$= \frac{m_{h_1^0}}{8\pi} (|\Delta_{(ab)L}|^2 + |\Delta_{(ab)R}|^2), \tag{36}$$

with the condition $m_{h_1^0}\gg m_{a,b}$. where $m_{a,b}$ are the masses of muon and tau, respectively. The on-shell conditions for external particles are $p_{1,2}^2=m_{a,b}^2$ and $p_{h_1^0}^2\equiv (p_1+p_2)^2=m_{h_1^0}^2$. The corresponding branching ratio is ${\rm Br}(h_1^0\to e_a e_b)=\Gamma(h_1^0\to e_a e_b)/\Gamma_{h_1^0}^{\rm total}$, where $\Gamma_{h_1^0}^{\rm total}\simeq 4.1\times 10^{-3}\,{\rm GeV}$ [21,26]. The $\Delta_{(ab)L,R}$ can be written as

$$\Delta_{(ab)L,R} = \sum_{i=1,5,7,8} \Delta_{(ab)L,R}^{(i)W} + \sum_{i=1}^{10} \Delta_{(ab)L,R}^{(i)Y}, \quad (37)$$

where the analytic forms of $\Delta_{(ab)L,R}^{(i)W}$ and $\Delta_{(ab)L,R}^{(i)Y}$ are shown in the Appendix. They can be calculated using the unitary gauge with the same techniques given in Refs. [6,22]. We have crosschecked this with FORM [27].

The divergence cancellation in the total amplitude (37) is proved analytically in the Appendix, based on the following strict equality:

$$U^{\nu}(\hat{M}^{\nu})^{2}U^{\nu\dagger} = (U^{\nu*}\hat{M}^{\nu}U^{\nu\dagger})^{*}U^{\nu*}\hat{M}^{\nu}U^{\nu\dagger} = M^{\nu*}M^{\nu}$$

$$= \begin{pmatrix} m_{D}^{*}m_{D}^{T} & 0 & m_{D}^{*}M_{R} \\ 0 & m_{D}^{\dagger}m_{D} + M_{R}^{*}M_{R}^{T} & M_{R}^{*}\mu_{X} \\ M_{R}^{\dagger}m_{D}^{T} & \mu_{X}^{*}M_{R}^{T} & M_{R}^{\dagger}M_{R} + \mu_{X}^{*}\mu_{X} \end{pmatrix}. \quad (38)$$

In the model under consideration, the divergent parts coming from the contributions of charged Higgs and heavy gauge bosons are related to both $(M^{\nu*}M^{\nu})_{(a+3)(b+3)}$ and $(M^{\nu*}M^{\nu})_{(a+6)(b+6)}$ $(a,b \leq 3)$, which are affected by heavy neutrino masses. The cancellation in the total divergent part

requires that the physical heavy neutrino masses and U^{ν} must be the exact values. Hence, approximate forms of the heavy neutrino masses and neutrino mixing matrix derived from the ISS mechanism cannot be applied. In contrast, we checked numerically that these formulas are safely used in the usual minimal ISS version extended directly from the SM, because the divergent parts are only involved with the elements $(M^{\nu*}M^{\nu})_{ab} = (m_D^*m_D^T)_{(ab)}$.

Many of the contributions listed in Eq. (37) are suppressed, and hence they can be ignored in our numerical computation. From now on, we just focus on the decay $h_1^0 \to \mu \tau$, and hence the simplified notations $\Delta_{L,R} \equiv \Delta_{(23)L,R}$ will be used. The decay $h_1^0 \to e \tau$ has similar properties, so we do not need to discuss it more explicitly. We can see that $|\frac{\Delta_L}{\Delta_R}| \simeq \mathcal{O}(\frac{m_\mu}{m_\tau})$. In addition, we prove in the Appendix that the following combinations are finite: $\Delta_{L,R}^{(1+5)W}$, $\Delta_{L,R}^{(7+8)W}$, $\Delta_{L,R}^{(4)YH_2}$, $\Delta_{L,R}^{(6+9+10)YH_2}$, $\Delta_{L,R}^{(4)YH_1}$, $\Delta_{L,R}^{(7+8)Y}$, and $(\Delta_{L,R}^{(1+2+3+5)Y} + \Delta_{L,R}^{(6+9+10)YH_1})$. With $m_{\mu,\tau} \ll m_W$, we have $B_1^{(1)} + B_1^{(2)}$, $B_1^{(2)} - B_0^{(2)} \simeq 0$, and hence $\Delta_{L,R}^{(7+8)W}$, $\Delta_{L,R}^{(7+8)Y} \simeq 0$. The two contributions $\Delta_{L,R}^{(4)YH_{1,2}^{\pm}}$ are also suppressed with a large $m_{H_2^{\pm}}$ for about a few TeV.

The four diagrams 4, 6, 9, and 10 in Fig. 1 include contributions from both charged Higgs bosons. They are not significantly affected by the $SU(3)_L$ scale m_Y , and thus they may enhance the partial decay widths of the LFVHDs if charged Higgs masses are small.

The regions of parameter space predicting large branching ratios for LFVHDs are affected strongly by the current experimental bound ${\rm Br}(\mu \to e \gamma) < 4.2 \times 10^{-13}$ [28]. A very good approximate formula for this decay rate in the limit m_u , $m_e \to 0$ is [11]

$$Br(\mu \to e\gamma) = \frac{12\pi^2}{G_F^2} |D_R|^2,$$
 (39)

where $G_F = g^2/(4\sqrt{2}m_W^2)$ and D_R is the one-loop contribution from charged gauge and Higgs boson mediations, $D_R = D_R^W + D_R^Y + D_R^{H_1^\pm} + D_R^{H_2^\pm}$. The analytic forms are

$$\begin{split} D_R^W &= -\frac{eg^2}{32\pi^2 m_W^2} \sum_{i=1}^9 U_{ai}^{\nu*} U_{bi}^{\nu} F(t_{iW}), \\ D_R^Y &= -\frac{eg^2}{32\pi^2 m_Y^2} \sum_{i=1}^9 U_{(a+3)i}^{\nu*} U_{(b+3)i}^{\nu} F(t_{iY}), \\ D_R^{H_{\pm}^{\pm}} &= -\frac{eg^2 f_k}{16\pi^2 m_W^2} \sum_{i=1}^9 \left[\frac{\lambda_{ai}^{L,k*} \lambda_{bi}^{L,k}}{m_{H_{\pm}^{\pm}}^2} \times \frac{1 - 6t_{ik} + 3t_{ik}^2 + 2t_{ik}^3 - 6t_{ik}^2 \ln(t_{i_k})}{12(t_{ik} - 1)^4} + \frac{m_{n_i} \lambda_{ai}^{L,k*} \lambda_{bi}^{\prime R,k}}{m_{H_{\pm}^{\pm}}^2} \times \frac{-1 + t_{ik}^2 - 2t_{ik} \ln(t_{ik})}{2(t_{ik} - 1)^3} \right], \end{split}$$
(40)

where

$$b = 2, \quad a = 1, \quad t_{iW} \equiv \frac{m_{n_i}^2}{m_W^2}, \quad t_{iY} \equiv \frac{m_{n_i}^2}{m_Y^2}, \quad t_{ik} \equiv \frac{m_{n_i}^2}{m_{H_k^{\pm}}^2},$$

$$f_1 \equiv \frac{1}{2}, \quad f_2 \equiv c_{\theta}^2, \quad \lambda_{bi}^{\prime R, 1} \equiv U_{(b+3)i}^{\nu}, \quad \lambda_{bi}^{\prime R, 2} \equiv U_{bi}^{\nu},$$

$$F(x) \equiv -\frac{10 - 43x + 78x^2 - 49x^3 + 4x^4 + 18x^3 \ln(x)}{12(x-1)^4}. \quad (41)$$

Because all charged Higgs bosons couple with heavy neutrinos through the Yukawa coupling matrix h_{ab}^{ν} , this matrix is strongly affected by the upper bound $\mathcal{O}(10^{-13})$ on $Br(\mu \to e\gamma)$. In fact, our numerical investigation shows that the allowed regions with light charged Higgs masses are very narrow. The previous investigation in Ref. [8] showed that the 331ISS model predicts a large Br($\mu \rightarrow e\gamma$), where the allowed regions discussed there were chosen such that $k \sim \mathcal{O}(10^3)$ and $M_R \leq 1$ TeV, implying that $z \sim \mathcal{O}(1)$ eV. We checked that our formulas are consistent with these results. In general, the allowed regions are very strict and satisfy one of the following conditions. First, the regions have a small z and large $|M_R|$ and $m_{H^{\pm}}$, implying $k \gg 1$, including those mainly discussed in Ref. [8]. Second, the regions allow for a large m_D and small k, but the strong destructive correlation between the two-loop contributions of charged gauge and Higgs bosons must happen. These regions were also considered in Ref. [8], but they were not given much attention. They are very interesting because they predict large branching ratios for LFVHDs and light particles such as new neutrinos and charged Higgs bosons, which could be found at the LHC and planned colliders [29,30]. Hence, our numerical investigation will focus on this case.

IV. NUMERICAL DISCUSSION ON LFVHDS

A. Setup parameters

To numerically investigate the LFVHDs of the SM-like Higgs boson, we will use the following well-known experimental parameters [21]: the mass of the W boson $m_W = 80.385$ GeV, the charged lepton masses $m_e = 5 \times 10^{-4}$ GeV, $m_\mu = 0.105$ GeV, and $m_\tau = 1.776$ GeV, the SM-like Higgs mass $m_{h_1^0} = 125.1$ GeV, and the gauge coupling of the $SU(2)_L$ symmetry $g \simeq 0.651$.

Combined with the discussion in Sec. II, the independent parameters are the heavy neutrino mass scale $M_R = \operatorname{diag}(M_R, M_R, M_R)$, the heavy gauge boson mass m_Y [considered as the $SU(3)_L$ breaking scale], the charged Higgs boson mass $m_{H_2^{\pm}}$, the characteristic scale of m_D defined as the parameter z, and the two Higgs self-couplings $\lambda_{1,12}$.

Other parameters can be calculated in terms of the above free ones, namely,

$$v_1 = v_2 = \frac{\sqrt{2}m_W}{g}, \qquad s_\theta = \frac{m_W}{m_Y\sqrt{2}}, \qquad w = \frac{2m_Y}{gc_\theta},$$

$$f = \frac{gc_\theta m_{H_2^{\pm}}^2}{4m_Y}, \qquad m_{H_1^{\pm}}^2 = \frac{m_{H_2^{\pm}}^2}{2}(t_\theta^2 + 1). \tag{42}$$

Apart from that, the mixing parameter α of the neutral *CP*-even Higgs was defined in Eq. (28). The Higgs self-coupling λ_2 is determined as [6]

$$\lambda_{2} = \frac{t_{\theta}^{2}}{2} \left(\frac{m_{h_{1}^{0}}^{2}}{v_{1}} - \frac{m_{H_{2}^{\pm}}^{2}}{2w^{2}} \right) + \frac{\left(\lambda_{12} - \frac{m_{H_{2}^{\pm}}^{2}}{2w^{2}} \right)^{2}}{4\lambda_{1} - \frac{m_{h_{1}^{0}}^{2}}{v_{1}^{2}}}.$$
 (43)

In the model under consideration with the quark sector given in Refs. [16,29], only the charged Higgs bosons H_2^\pm couple with all SM leptons and quarks. They have been investigated at the LHC in the direct production $pp \to t(b)H^\pm$, which then decay into two final fermion states [31]. But the specific constraints on them in the framework of the 3-3-1 models have not been mentioned yet, to the best of our knowledge. Instead, the lower bounds on their masses have been discussed recently based on recent data of neutral meson mixing $B_0 - \bar{B}_0$, where a reasonable lower bound of $m_{H_2^\pm} \ge 480$ GeV was concerned [29].

The values of $\lambda_{1,2,12}$ must satisfy theoretical conditions of unitarity and the Higgs potential must be bounded from below, as mentioned in Ref. [6]. The heavy charged gauge boson mass m_Y is related to the recent lower constraint of neutral gauge boson Z' in this model.

For the above reasons, the default values of the free parameters chosen for our numerical investigation are as follows. Without loss of generality, the Higgs self-couplings are fixed as $\lambda_1=1$, $\lambda_{12}=-1$, which also guarantee that all couplings of the SM-like Higgs boson approach the SM limit when $t_\theta \to 0$. The default value $m_Y=4.5$ TeV satisfies all recent constraints [29,32]. The parameter z will be considered in the range of the perturbative limit $z<2\sqrt{\pi}\times v_1\simeq 617$ GeV; in particular, we will fix z=50,200,400,500, and 600 [GeV]. Finally, the charged Higgs mass $m_{H_2^\pm}$ will be investigated mainly in the range of 300 to 5×10^4 GeV, where large values of LFVHDs may appear.

B. Numerical results

First, we reproduce the regions mentioned in Ref. [8], where M_R was chosen to be from hundreds of GeV to 1 TeV, and the scale of m_D (namely, z) was a few GeV, corresponding to $k\gg 1$. As a result, the respective regions of parameter space always satisfy the experimental bound on $\text{Br}(\mu\to e\gamma)$ with large enough $m_{H_2^\pm}$. These regions are shown in Fig. 2 with fixed z=1, 5, 10, 100, and 500 GeV. All allowed regions (i.e., those that satisfy the upper bound $\text{Br}(\mu\to e\gamma) < 4.2 \times 10^{-13}$) give a small $\text{Br}(h_1^0\to \mu\tau) < \mathcal{O}(10^{-9})$. In general, for larger k we checked numerically that the values of the branching ratio of LFVHDs will decrease significantly, and hence we will not discuss this further.

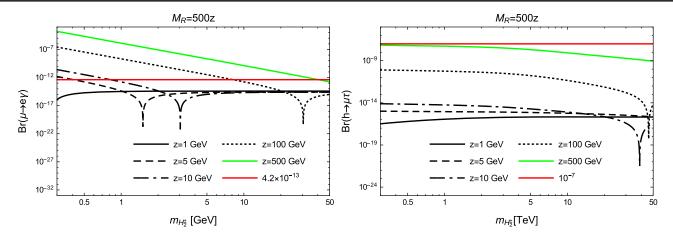


FIG. 2. Br($\mu \to e\gamma$) (left) and Br($h_1^0 \to \mu\tau$) (right) as functions of $m_{H_2^{\pm}}$ with k = 500.

With small values of k=5.5 and 9, the dependence of both $\text{Br}(\mu \to e \gamma)$ and $\text{Br}(h \to \mu \tau)$ on $m_{H_2^\pm}$ with fixed z are shown in Fig. 3. Most regions of the parameter space are ruled out by the bound on $\text{Br}(\mu \to e \gamma)$, except for narrow parts where particular contributions from charged Higgs and gauge bosons are destructive. This interesting property of the 331ISS model was indicated previously in Ref. [8].

Furthermore, it predicts allowed regions that give a large ${\rm Br}(h_1^0 \to \mu \tau)$. In particular, the largest values can reach $\mathcal{O}(10^{-4})$ when k=5.5 and z=600 GeV, which is very close to the perturbative limit. In general, the illustrations in two Figs. 2 and 3 suggest that this branching ratio is enhanced significantly for smaller k and larger z, but changes slowly with the change of large $m_{H_3^\pm}$. In contrast,

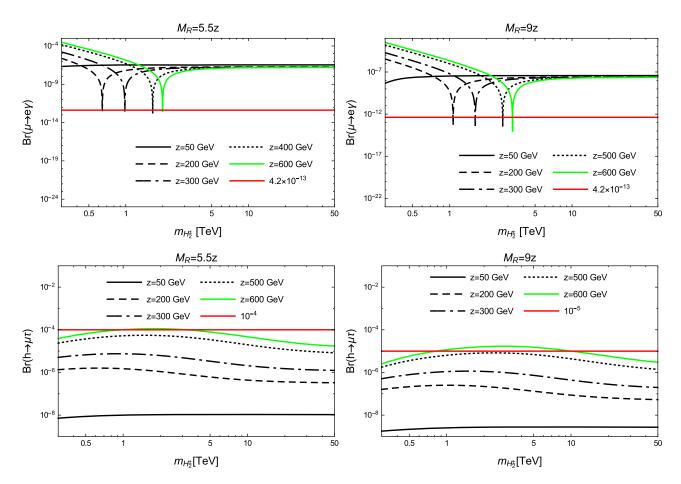


FIG. 3. Br($\mu \to e\gamma$) (upper) and Br($h_1^0 \to \mu\tau$) (lower) as functions of $m_{H_2^{\pm}}$ with k = 5.5 (left) and k = 9 (right).

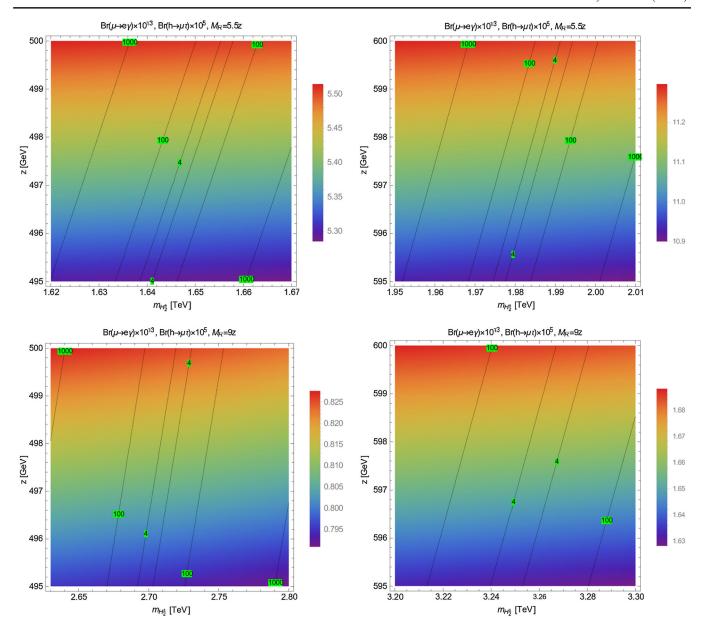


FIG. 4. Density plots of $Br(h_1^0 \to \mu \tau)$ and contour plots of $Br(\mu \to e \gamma)$ (black curves) as functions of $m_{H_2^{\pm}}$ and z, with k = 5.5 (upper) and k = 9 (lower).

a small $m_{H_2^\pm}$ plays a very important role in creating allowed regions that predict a large LFVHD. Br $(\mu \to e \gamma)$ does not depend on $m_{H_2^\pm}$ when it is large enough. Furthermore, the branching ratio decreases with increasing k and it will go below the experimental bound if k is large enough.

The allowed regions in Fig. 3 are shown more explicitly in Fig. 4, corresponding to k=5.5 and k=9. Only regions that give a large ${\rm Br}(h_1^0\to\mu\tau)$ are mentioned. They are bounded between two black curves representing the constant value of ${\rm Br}(\mu\to e\gamma)\times 10^{13}=4$. Clearly, ${\rm Br}(h_1^0\to\mu\tau)$ is sensitive to z and k, while it changes slowly with changing values of $m_{H_2^\pm}$. In contrast, the suppressed ${\rm Br}(\mu\to e\gamma)$ allows narrow regions of the parameter space

where some particular relation between $m_{H_2^{\pm}}$ and k and z is realized. Hence, if these two decay channel are discovered by experiments, depending on their specific values, a relation between heavy neutrino and charged Higgs masses can be determined from the 331ISS framework.

To understand how $\operatorname{Br}(h_1^0 \to \mu \tau)$ depends on the $SU(3)_L$ breaking scale defined by m_Y in this work, four allowed regions corresponding to the four fixed values $m_Y = 3, 4, 5$, and 6 TeV are illustrated in Fig. 5. It can be seen that the branching ratio of LFVHD depends weakly on m_Y , namely, it decreases slowly with increasing m_Y . Hence, studies of LFV decays will give useful information about heavy neutrinos and charged Higgs bosons besides the phenomenology arising from heavy gauge bosons discussed in many

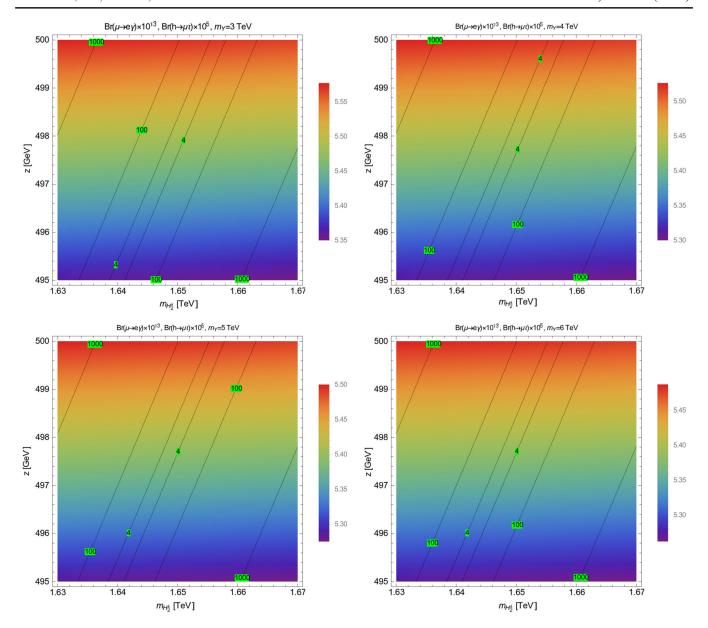


FIG. 5. Density plots of $Br(h_1^0 \to \mu \tau)$ and contour plots of $Br(\mu \to e \gamma)$ (black curves) as functions of $m_{H_2^{\pm}}$ and z, with k = 5.5, z around 500 GeV, and different m_Y .

earlier works. More interestingly, this may happen at large $SU(3)_L$ scales which the LHC cannot detect at present.

V. CONCLUSION

The 331ISS models seem to be the most interesting among the well-known 331 models because of their rich phenomenology, as indicated in many recent works. This work addressed a more attractive property, namely, the LFVHDs of the SM-like Higgs boson which are being investigated at the LHC. Assuming the absence of the tree-level decays $h_1^0 \rightarrow e_a e_b$ and $e_j \rightarrow e_i \gamma$ (j > i), the analytical formulas at the one-loop level to calculate these decay rates in the 331ISS model have been introduced. The divergent cancellation in the total decay amplitudes of $h_1^0 \rightarrow e_a e_b$ was

shown explicitly. From the numerical investigation, we have indicated that the ${\rm Br}(h_1^0 \to \mu \tau)$ predicted by the 331ISS model can reach large values of $\mathcal{O}(10^{-5})$. They are even very close to 10^{-4} , for example, in the special case with k=5.5 and $z\simeq 600$ GeV, which is close to the perturbative limit of the lepton Yukawa couplings. This value is larger than that predicted by the simplest ISS version extended directly from the SM [17]. New charged Higgs bosons may give large contributions to both of the decay rates ${\rm Br}(\mu \to e \gamma)$ and ${\rm Br}(h_1^0 \to \mu \tau)$, leading to either constructive or destructive correlations with those from the charged gauge bosons. As a by-product, the recent experimental bound on ${\rm Br}(\mu \to e \gamma)$ rules out most of the regions of parameter space with small k and large z, except the

narrow regions arising from the destructive correlations between contributions of charged Higgs and gauge bosons. We have shown numerically that only these regions give a large $Br(h_1^0 \to \mu \tau) > 10^{-5}$ when 400 GeV < z < 600 GeV and $k \le 9$ in the case where the Majorana mass matrix M_R is proportional to the identity one. Furthermore, these large values of $Br(h_1^0 \to \mu \tau)$ depend weakly on the masses of the heavy charged gauge bosons, but they require the heavy neutrino mass scale M_R and $m_{H_2^{\pm}}$ to be a few TeV, which can be detected at current colliders. Besides, $Br(h_1^0 \to \tau e)$ has the same result. In conclusion, large branching ratios of the LFV processes like $h_1^0 \to \mu \tau, e \tau$ will support the 331ISS model and may rule out the original 331RHN model containing only very light exotic neutrinos. Additionally, many properties of heavy neutrinos and charged Higgs bosons in the 331ISS framework may be determined independently at the $SU(3)_L$ scale.

ACKNOWLEDGMENTS

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APPENDIX: FORM FACTORS OF LFVHDs IN THE UNITARY GAUGE

In this appendix we list all analytic formulas of oneloop contributions to LFVHDs defined in Eq. (37). They are written in terms of Passarino-Veltman functions that were defined thoroughly in Refs. [6,22]. Using the notations $D_0 = k^2 - M_0^2 + i\delta$, $D_1 = (k - p_1)^2 - M_1^2 + i\delta$, and $D_2 = (k + p_2)^2 - M_2^2 + i\delta$, where δ is an infinitesimal positive real quantity, the one-loop integrals and Passarino-Veltman functions needed in this work are

$$\begin{split} B_{0,\mu}^{(i)} &\equiv \frac{(2\pi\mu)^{4-D}}{i\pi^2} \int \frac{d^D k\{1,k_\mu\}}{D_0 D_i} \,, \qquad B_0^{(12)} &\equiv \frac{(2\pi\mu)^{4-D}}{i\pi^2} \int \frac{d^D k}{D_1 D_2} \,, \\ C_{0,\mu} &\equiv C_{0,\mu}(M_0,M_1,M_2) = \frac{1}{i\pi^2} \int \frac{d^4 k\{1,k_\mu\}}{D_0 D_1 D_2} \,, \qquad B_\mu^{(i)} &= B_1^{(i)} p_{i\mu}, \qquad C_\mu = C_1 p_{1\mu} + C_2 p_{2\mu}, \end{split}$$

where i=1,2. In addition, $D=4-2\epsilon \le 4$ is the dimension of the integral, M_0,M_1,M_2 are masses of virtual particles in the loop, and μ is an arbitrary mass parameter introduced via dimensional regularization [33]. The external momenta of the final leptons shown in Fig. 1 satisfy $p_1^2 = m_a^2$, $p_2^2 = m_b^2$ and $(p_1 + p_2)^2 = m_{h_1^0}^2$, where $m_{h_1^0}$ is the SM-like Higgs boson mass, and

 $m_{a,b}$ are lepton masses. In the limit $m_{a,b} \simeq 0$, the analytic formulas for $B_{0,1}^{(i)}, B_0^{(12)}, C_0$, and $C_{1,2}$ were shown in Refs. [6,22,34], and hence we will not repeat them here. These functions are used for our numerical investigation. We stress that they were checked numerically to be well consistent with the exact results computed using LOOPTOOLS [35], as reported in Ref. [36]. The analytic expressions for $\Delta_{L,R}^{(i)W} \equiv \Delta_{(ab)L,R}^{(i)W}$, where *i* implies the diagram (*i*) in Fig. 1, are

$$\begin{split} \Delta_L^{(1)W} &= \frac{g^3 c_a m_a}{64 \pi^2 m_W^3} \sum_{i=1}^9 U_{ai}^{\nu*} U_{bi}^{\nu} \{ m_{n_i}^2 (B_1^{(1)} - B_0^{(1)} - B_0^{(2)}) - m_b^2 B_1^{(2)} + (2 m_W^2 + m_{h_i}^2) m_{n_i}^2 C_0 \\ &- [2 m_W^2 (2 m_W^2 + m_{n_i}^2 + m_a^2 - m_b^2) + m_{n_i}^2 m_{h_i}^2] C_1 + [2 m_W^2 (m_a^2 - m_{h_i}^2) + m_b^2 m_{h_i}^2] C_2 \}, \\ \Delta_R^{(1)W} &= \frac{g^3 c_a m_b}{64 \pi^2 m_W^3} \sum_{i=1}^9 U_{ai}^{\nu*} U_{bi}^{\nu} \{ -m_{n_i}^2 (B_1^{(2)} + B_0^{(1)} + B_0^{(2)}) + m_a^2 B_1^{(1)} + (2 m_W^2 + m_{h_i}^2) m_{n_i}^2 C_0 \\ &- [2 m_W^2 (m_b^2 - m_h^2) + m_a^2 m_{h_i}^2] C_1 + [2 m_W^2 (2 m_W^2 + m_{n_i}^2 - m_a^2 + m_b^2) + m_{n_i}^2 m_{h_i}^2] C_2 \}, \\ \Delta_L^{(5)W} &= \frac{g^3 c_a m_a}{64 \pi^2 m_W^3} \sum_{i,j=1}^9 U_{ai}^{\nu*} U_{bj}^{\nu} \{ \lambda_{ij}^0 m_{n_j} [B_0^{(12)} - m_W^2 C_0 + (2 m_W^2 + m_{n_i}^2 - m_a^2) C_1] + \lambda_{ij}^0 m_{n_i} [B_1^{(1)} + (2 m_W^2 + m_{n_j}^2 - m_b^2) C_1] \}, \\ \Delta_R^{(5)W} &= \frac{g^3 c_a m_b}{64 \pi^2 m_W^3} \sum_{i=1}^9 U_{ai}^{\nu*} U_{bj}^{\nu} \{ \lambda_{ij}^0 m_{n_i} [B_0^{(12)} - m_W^2 C_0 - (2 m_W^2 + m_{n_j}^2 - m_b^2) C_2] - \lambda_{ij}^0 m_{n_j} [B_1^{(2)} + (2 m_W^2 + m_{n_i}^2 - m_a^2) C_2] \}, \\ \Delta_L^{(7+8)W} &= \frac{g^3 m_a m_b^2 c_a}{64 \pi^2 m_W^3 (m_a^2 - m_b^2)} \sum_{i=1}^9 U_{ai}^{\nu*} U_{bi}^{\nu} [2 m_{n_i}^2 (B_0^{(1)} - B_0^{(2)}) - (2 m_W^2 + m_{n_i}^2) (B_1^{(1)} + B_1^{(2)}) - m_a^2 B_1^{(1)} - m_b^2 B_2^{(1)}], \\ \Delta_R^{(7+8)W} &= \frac{m_a}{m_b} \Delta_L^{(7+8)W}. \end{split}$$

Defining $\Delta_{L,R}^{(i)Y} = \Delta_{(ab)L,R}^{(i)YH_1^{\pm}} + \Delta_{(ab)L,R}^{(i)YH_2^{\pm}}$ with i = 4, 6, 9, 10, the analytic expressions for $\Delta_{L,R}^{(i)Y} \equiv \Delta_{(ab)L,R}^{(i)Y}$ are

$$\begin{split} &\Delta_{L}^{(1)Y} = -\frac{g^{2}m_{\alpha}(\sqrt{2}s_{\alpha}c_{\alpha}-c_{\alpha}s_{\alpha})}{644\sqrt{2}x^{2}m_{N}^{2}} \sum_{j=1}^{j} U_{(e-3)j}^{j} U_{(e-3)j}^{j} \{m_{\alpha}^{2}(m_{\alpha}^{2}(H_{1}^{2}-B_{0}^{1})-m_{\delta}^{2}B_{1}^{2}) \\ &+ (2m_{V}^{2}+m_{\phi}^{2})m_{\kappa}^{2}C_{0} - [2m_{V}^{2}(c_{W}^{2}+m_{\phi}^{2}+m_{\kappa}^{2}+m_{\delta}^{2})+m_{\kappa}^{2}m_{\delta}^{2}] + m_{\delta}^{2}m_{\delta}^{2}]_{0} + k_{\delta}^{2}m_{\delta}^{2}C_{0} + m_{\delta}^{2}m_{\delta}^{2}C_{0} + m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}C_{0} + [2m_{V}^{2}(m_{\sigma}^{2}-m_{\phi}^{2})+m_{\delta}^{2}m_{\delta}^{2}]_{0}^{2}C_{0} + m_{\delta}^{2}m_{\delta}^{2}C_{0}^{2}) \\ &+ (2m_{V}^{2}+m_{\phi}^{2})m_{\kappa}^{2}(C_{0}-C_{0}^{2}m_{\delta}^{2}) - 2m_{\phi}^{2}m_{\delta}^{2}) + m_{\delta}^{2}m_{\delta}^{2}[C_{1} + [2m_{V}^{2}(2m_{V}^{2}+m_{\phi}^{2}-m_{\phi}^{2}+m_{\delta}^{2})+m_{\delta}^{2}m_{\delta}^{2}]C_{2} \}, \\ &+ (2m_{V}^{2}+m_{\phi}^{2})m_{\kappa}^{2}(C_{0}-V_{0}^{2}+S_{\delta}^{2}) - 2m_{\phi}^{2}m_{\delta}^{2}) + m_{\delta}^{2}m_{\delta}^{2}[C_{1} + [2m_{V}^{2}(2m_{V}^{2}+m_{\phi}^{2}-m_{\phi}^{2}+m_{\delta}^{2})+m_{\phi}^{2}m_{\delta}^{2}]C_{2} \}, \\ &+ (2m_{V}^{2}+m_{\phi}^{2})m_{\kappa}^{2}C_{0}^{2}(C_{0}^{2}+V_{0}^{2}+S_{\delta}^{2}a_{\delta}^{2}) - 2m_{\phi}^{2}m_{\delta}^{2}M_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}[C_{1} + [2m_{V}^{2}(2m_{V}^{2}+m_{\phi}^{2}-m_{\phi}^{2}+m_{\phi}^{2}+m_{\phi}^{2}]C_{2}], \\ &+ k_{\delta}^{2}M_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}C_{\delta}^{2}(C_{\delta}^{2}+V_{0}^{2}+S_{\delta}^{2}a_{\delta}^{2}) - 2m_{\phi}^{2}m_{\delta}^{2}M_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}(C_{\delta}^{2}) + 2m_{\phi}^{2}m_{\delta}^{2}C_{2}] \}, \\ &\Delta_{E}^{(1)Y} = \frac{g^{2}c_{\phi}(c_{\phi}c_{\phi}^{2}+V_{0}^{2}+S_{\phi}^{2}a_{\delta}^{2})}{64\pi^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}} - m_{\phi}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}C_{2}] \}, \\ &\Delta_{E}^{(1)Y} = \frac{g^{2}c_{\phi}(c_{\phi}c_{\phi}^{2}+V_{0}^{2}+S_{\phi}^{2}a_{\delta}^{2})}{64\pi^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}} - m_{\phi}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}C_{2}] + m_{\phi}^{2}m_{\delta}^{2}C_{2}] \}, \\ &\Delta_{E}^{(1)Y} = \frac{g^{2}c_{\phi}(c_{\phi}c_{\phi}^{2}+V_{0}^{2}+S_{\phi}^{2}a_{\delta}^{2})}{64\pi^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}} - m_{\phi}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}C_{2}] + m_{\phi}^{2}m_{\delta}^{2}C_{2}] \}, \\ &\Delta_{E}^{(1)Y} = \frac{g^{2}c_{\phi}(c_{\phi}c_{\phi}^{2}+V_{0}^{2}+S_{\phi}^{2}a_{\delta}^{2})}{64\pi^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}m_{\delta}^{2}} - m_{\phi}^{2}m_{\delta}^{2}m_{\delta}^{2}C_{2$$

$$\begin{split} \Delta_L^{(7+8)Y} &= \frac{g^3 m_a m_b^2 c_\alpha}{64 \pi^2 m_W m_Y^2 (m_a^2 - m_b^2)} \sum_{i=1}^9 U_{(a+3)i}^{\nu*} U_{(b+3)i}^{\nu} [2 m_{n_i}^2 (B_0^{(1)} - B_0^{(2)}) - (2 m_Y^2 + m_{n_i}^2) (B_1^{(1)} + B_1^{(2)}) - m_a^2 B_1^{(1)} - m_b^2 B_1^{(2)}], \\ \Delta_R^{(7+8)Y} &= \frac{m_a}{m_b} \Delta_L^{(7+8)Y}, \\ \Delta_L^{(9+10)YH_k^{\pm}} &= -\frac{g^3 c_\alpha f_k}{32 \pi^2 m_W^3 (m_a^2 - m_b^2)} \sum_{i=1}^9 [m_a m_b m_{n_i} \lambda_{ai}^{L,k*} \lambda_{bi}^{R,k} (B_0^{(1)} - B_0^{(2)}) + m_{n_i} \lambda_{ai}^{R,k*} \lambda_{bi}^{L,k} (m_b^2 B_0^{(1)} - m_a^2 B_0^{(2)}) \\ &+ m_a m_b (\lambda_{ai}^{L,k*} \lambda_{bi}^{L,k} m_b + \lambda_{ai}^{R,k*} \lambda_{bi}^{R,k} m_a) (B_1^{(1)} + B_1^{(2)})], \\ \Delta_R^{(9+10)YH_k^{\pm}} &= -\frac{g^3 c_\alpha f_k}{32 \pi^2 m_W^3 (m_a^2 - m_b^2)} \sum_{i=1}^9 [m_a m_b m_{n_i} \lambda_{ai}^{R,k*} \lambda_{bi}^{L,k} (B_0^{(1)} - B_0^{(2)}) + m_{n_i} \lambda_{ai}^{L,k*} \lambda_{bi}^{R,k} (m_b^2 B_0^{(1)} - m_a^2 B_0^{(2)}) \\ &+ m_a m_b (\lambda_{ai}^{R,k*} \lambda_{bi}^{R,k} m_b + \lambda_{ai}^{L,k*} \lambda_{bi}^{L,k} m_a) (B_1^{(1)} + B_1^{(2)})], \end{split} \tag{A2}$$

where $f_1 = c_\theta^2$ and $f_2 = 1/2$. The details to derive the expressions in Eq. (A2) are the same as those shown in Refs. [6,22], and hence we do not present them in this work. We note that the scalar functions $\Delta_{L,R}^{(1)W}$ and $\Delta_{L,R}^{(1,2,3)Y}$ include parts that do not depend on m_{n_i} , and therefore they vanish because of the Glashow-Iliopoulos-Maiani mechanism. They are ignored in Eqs. (A1) and (A2).

The divergent cancellation in the total $\Delta_{L,R}$ is shown as follows. The divergent parts only contain B functions: $\operatorname{div} B_0^{(1)} = \operatorname{div} B_0^{(2)} = \operatorname{div} B_0^{(12)} = 2 \operatorname{div} B_1^{(1)} = -2 \operatorname{div} B_1^{(2)} = \Delta_{\epsilon}$. Ignoring the common factor of $g^3/(64\pi^2 m_W^3)$ and using $1/m_Y = \sqrt{2} s_\theta/m_W$, the divergent parts of Δ_L derived from Eq. (A2) are

$$\begin{split} \operatorname{div}[\Delta_{L}^{(1)W}] &= m_{a} \Delta_{\epsilon} \times \left(-\frac{3c_{a}}{2}\right) \sum_{i=1}^{9} U_{ai}^{\nu s} U_{bi}^{\nu} m_{n_{i}}^{2}, \\ \operatorname{div}[\Delta_{L}^{(S)W}] &= m_{a} \Delta_{\epsilon} \times c_{a} \sum_{i,j=1}^{9} U_{ai}^{\nu s} U_{bj}^{\nu} \left(\lambda_{ij}^{0*} m_{n_{j}} + \frac{1}{2} \lambda_{ij}^{0} m_{n_{i}}\right), \\ \operatorname{div}[\Delta_{L}^{(7+8)W}] &= \operatorname{div}[\Delta_{L}^{(4)Y}] &= \operatorname{div}[\Delta_{L}^{(7+8)Y}] = 0, \\ \operatorname{div}[\Delta_{L}^{(1)Y}] &= m_{a} \Delta_{\epsilon} \times 3s_{\theta}^{3} (\sqrt{2}s_{a}c_{\theta} - c_{a}s_{\theta}) \sum_{i=1}^{9} U_{(a+3)i}^{\nu s} U_{(b+3)i}^{\nu} m_{n_{i}}^{2}, \\ \operatorname{div}[\Delta_{L}^{(2)Y}] &= m_{a} \Delta_{\epsilon} \times s_{\theta}^{2} c_{\theta} (c_{a}c_{\theta} + \sqrt{2}s_{a}s_{\theta}) \sum_{i=1}^{9} U_{(a+3)i}^{\nu s} \lambda_{bi}^{L_{1}} m_{n_{i}}, \\ \operatorname{div}[\Delta_{L}^{(3)Y}] &= m_{a} \Delta_{\epsilon} \times \left[-2s_{\theta}^{2} c_{\theta} (c_{a}c_{\theta} + \sqrt{2}s_{a}s_{\theta})\right] \sum_{i=1}^{9} U_{(a+3)i}^{\nu s} U_{(b+3)i}^{\nu} m_{n_{i}}^{2}, \\ \operatorname{div}[\Delta_{L}^{(5)Y}] &= m_{a} \Delta_{\epsilon} \times 2s_{\theta}^{2} c_{a} \sum_{i,j=1}^{9} U_{(a+3)i}^{\nu s} U_{(b+3)j}^{\nu} \left(\lambda_{ij}^{0*} m_{n_{j}} + \frac{1}{2} \lambda_{ij}^{0} m_{n_{i}}\right), \\ \operatorname{div}[\Delta_{L}^{(6)YH_{1}^{\pm}}] &= m_{a} \Delta_{\epsilon} \times \left(-2c_{a}c_{\theta}^{2}\right) \sum_{i,j=1}^{9} U_{(a+3)i}^{\nu s} \lambda_{bj}^{0*} \lambda_{bj}^{L,1}, \\ \operatorname{div}[\Delta_{L}^{(6)YH_{2}^{\pm}}] &= m_{a} \Delta_{\epsilon} \times (2c_{a}c_{\theta}^{2}) \sum_{i,j=1}^{9} U_{ai}^{\nu s} \lambda_{ij}^{0*} \lambda_{bi}^{L,1} m_{n_{i}}, \\ \operatorname{div}[\Delta_{L}^{(9+10)YH_{1}^{\pm}}] &= m_{a} \Delta_{\epsilon} \times c_{a} \sum_{i=1}^{9} U_{ai}^{\nu s} \lambda_{bi}^{L,2} m_{n_{i}}, \\ \operatorname{div}[\Delta_{L}^{(9+10)YH_{2}^{\pm}}] &= m_{a} \Delta_{\epsilon} \times c_{a} \sum_{i=1}^{9} U_{ai}^{\nu s} \lambda_{bi}^{L,2} m_{n_{i}}. \end{cases} \tag{A3}$$

Using the equalities $M^{\nu} = U^{\nu*} \hat{M}^{\nu} U^{\nu\dagger}$ and Eq. (38), we can prove that

$$\begin{split} \operatorname{div}[\Delta_{L,R}^{(1)W}] &\sim \sum_{i=1}^{9} U_{ai}^{\nu*} U_{bi}^{\nu} m_{n_{i}}^{2} = \left[U^{\nu} (\hat{M}^{\nu})^{2} U^{\nu^{\dagger}} \right]_{ba} = (m_{D}^{*} m_{D}^{T})_{ba} = (m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(5)W}] &\sim \sum_{i,j=1}^{9} U_{ai}^{\nu*} U_{bj}^{\nu} \lambda_{ij}^{0*} m_{n_{j}}, \qquad \sum_{i,j=1}^{9} U_{ai}^{\nu*} U_{bj}^{\nu} \lambda_{ij}^{0} m_{n_{i}} = (m_{D}^{*} m_{D}^{T})_{ba} = (m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(1,3)Y}] &\sim \sum_{i=1}^{9} U_{(a+3)i}^{\nu*} U_{(b+3)i}^{\nu} m_{n_{i}}^{2} = \left[U^{\nu} (\hat{M}^{\nu})^{2} U^{\nu^{\dagger}} \right]_{(b+3)(a+3)} = (m_{D}^{\dagger} m_{D} + M_{R}^{*} M_{R}^{T})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(2)Y,(9+10)YH_{1}^{\dagger}}] &\sim \sum_{i=1}^{9} U_{(a+3)i}^{\nu*} \lambda_{bi}^{0*} m_{n_{i}} = (m_{D}^{*} m_{D})_{ba} + t_{\theta}^{2} (M_{R}^{*} M_{R}^{T})_{ba} = -(m_{D}^{\dagger} m_{D})_{ba} + t_{\theta}^{2} (M_{R}^{*} M_{R}^{T})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(5)Y}] &\sim \sum_{i,j=1}^{9} U_{(a+3)i}^{\nu*} U_{(b+3)j}^{0} \lambda_{ij}^{0*} m_{n_{i}}, \sum_{i,j=1}^{9} U_{(a+3)i}^{\nu*} U_{(b+3)j}^{\nu} \lambda_{ij}^{0} m_{n_{i}} \sim (m_{D}^{\dagger} m_{D})_{ba} - \sqrt{2} t_{a} t_{\theta} (M_{R}^{*} M_{R}^{T})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(6)YH_{1}^{\dagger}}] &\sim \sum_{i,j=1}^{9} U_{(a+3)i}^{\nu*} \lambda_{ij}^{0*} \lambda_{bj}^{L,1} = (m_{D}^{*} m_{D})_{ba} - \sqrt{2} t_{a} t_{\theta}^{3} (M_{R}^{*} M_{R}^{T})_{ba} = -(m_{D}^{\dagger} m_{D})_{ba} - \sqrt{2} t_{a} t_{\theta}^{3} (M_{R}^{*} M_{R}^{T})_{ba} = -(m_{D}^{\dagger} m_{D})_{ba} - \sqrt{2} t_{a} t_{\theta}^{3} (M_{R}^{*} M_{R}^{T})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(6)YH_{2}^{\dagger}}] &\sim \sum_{i,j=1}^{9} U_{ai}^{\nu*} \lambda_{ij}^{0*} \lambda_{bj}^{L,2} = -(m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(9+10)YH_{2}^{\dagger}}] &\sim \sum_{i=1}^{9} U_{ai}^{\nu*} \lambda_{ij}^{0*} \lambda_{bi}^{L,2} = -(m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(9+10)YH_{2}^{\dagger}}] &\sim \sum_{i=1}^{9} U_{ai}^{\nu*} \lambda_{bi}^{1} \lambda_{bi}^{L,2} m_{n_{i}} = -(m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(9+10)YH_{2}^{\dagger}}] &\sim \sum_{i=1}^{9} U_{ai}^{\nu*} \lambda_{bi}^{L,2} m_{n_{i}} = -(m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(9+10)YH_{2}^{\dagger}}] &\sim \sum_{i=1}^{9} U_{ai}^{\nu*} \lambda_{bi}^{L,2} m_{n_{i}} = -(m_{D}^{\dagger} m_{D})_{ba}, \\ \operatorname{div}[\Delta_{L,R}^{(9+10)YH_{2}^{\dagger}}] &\sim \sum_{i=1}^{9} U_{ai}^{\nu*} \lambda_{bi}^{L,2} m_{n_{i}} = -(m_{D}^{\dagger} m_{D})_{ba}, \\$$

where we have used the antisymmetric property of m_D : $m_D^T = -m_D$. From this, it can be seen that $\operatorname{div}[\Delta_L^{(1)W}] + \operatorname{div}[\Delta_L^{(5)W}] = \operatorname{div}[\Delta_L^{(6)YH_2^{\pm}}] + \operatorname{div}[\Delta_L^{(9+10)YH_2^{\pm}}] = 0$. The sum of the remaining divergent parts is

$$\begin{split} \operatorname{div}[\Delta_{L}^{(1+2+3+5)Y} + \Delta_{L}^{(6+9+10)YH_{1}^{\pm}}] &\sim (m_{D}^{\dagger}m_{D})_{ba}\{\sqrt{2}s_{\alpha}s_{\theta}^{2}c_{\theta}(3-1-2) + c_{\alpha}[s_{\theta}^{2}(-3s_{\theta}^{2}-c_{\theta}^{2}-2c_{\theta}^{2}+3) + 2s_{\theta}^{2}-2s_{\theta}^{2}]\} \\ &\quad + (M_{R}^{*}M_{R}^{T})_{ba}[\sqrt{2}s_{\alpha}\frac{s_{\theta}^{2}}{c_{\theta}}(3c_{\theta}^{2}+s_{\theta}^{2}-2c_{\theta}^{2}-3+2) + c_{\alpha}s_{\theta}^{2}(-3s_{\theta}^{2}+s_{\theta}^{2}-2c_{\theta}^{2}+2)] \\ &= 0. \end{split} \tag{A5}$$

Finally, the proof of the divergent cancellation in Δ_R is exactly the same as that in Δ_L .

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