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An Invitation to Quantum Field Theory

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Contents

1	Why Do We Need Quantum Field Theory After All?	1
1.1	Relativistic Quantum Mechanics	1
1.2	The Klein Paradox	4
1.3	From Wave Functions to Quantum Fields	6
	References	8
2	From Classical to Quantum Fields	11
2.1	Particles and Quantum Fields	11
2.2	Canonical Quantization	17
2.3	The Casimir Effect	22
2.4	Path Integrals	24
2.5	The Semiclassical Limit	28
	References	32
3	Theories and Lagrangians I: Matter Fields	33
3.1	Representations of the Lorentz Group	33
3.2	Weyl Spinors	36
3.3	Dirac Spinors	38
4	Theories and Lagrangians II: Introducing Gauge Fields	47
4.1	Classical Gauge Fields	47
4.2	Quantization of the Electromagnetic Field	55
4.3	Coupling Gauge Fields to Matter	56
4.4	Nonabelian Gauge Theories	58
4.5	Understanding Gauge Symmetry	61
4.6	Gauge Fields and Path Integrals	64
4.7	The Structure of the Gauge Theory Vacuum	68
4.8	Instantons in Gauge Theories	75
	References	78

5	Theories and Lagrangians III: The Standard Model	81
5.1	Fundamental Interactions	81
5.2	Leptons and Quarks	85
5.3	Quantum Chromodynamics	90
5.4	The Electroweak Theory	92
5.5	Closing Remarks: Particle Masses in the Standard Model	98
	References	99
6	Towards Computational Rules: Feynman Diagrams	101
6.1	Cross Sections and S-Matrix Amplitudes	101
6.2	From Green's Functions to Scattering Amplitudes	109
6.3	Feynman Rules	110
6.4	An Example: Compton Scattering at Low Energies	116
6.5	Polarization of the Cosmic Microwave Background Radiation	121
	References	126
7	Symmetries I: Continuous Symmetries	127
7.1	Noether's Theorem	127
7.2	Quantum Mechanical Realizations of Symmetries	131
7.3	Some Applications of Goldstone's Theorem	136
7.4	The Brout–Englert–Higgs Mechanism	141
	References	144
8	Renormalization	145
8.1	Removing Infinities	145
8.2	The Beta-Function and Asymptotic Freedom	150
8.3	A Look at the Systematics of Renormalization	155
8.4	Renormalization in Statistical Mechanics	162
8.5	The Renormalization Group in Quantum Field Theory	167
	References	172
9	Anomalies	175
9.1	A Toy Model for the Axial Anomaly	175
9.2	The Triangle Diagram	181
9.3	Chiral Symmetry in QCD	183
9.4	Gauge Anomalies	189
	References	192
10	The Origin of Mass	193
10.1	The Masses in the Standard Model	193
10.2	Quark Masses	202
10.3	Λ_{QCD} and the Hadron Masses	204
	References	208

11	Symmetries II: Discrete Symmetries	209
11.1	Discrete Symmetries in Classical Mechanics and Field Theory	209
11.2	Parity and Charge Conjugation in Quantum Field Theory	213
11.3	Majorana Spinors	215
11.4	Time Reversal	216
11.5	CP Symmetry and CP Violation	219
11.6	The CPT Theorem	222
11.7	Spin and Statistics	229
	References	230
12	Effective Field Theories and Naturalness	231
12.1	Energy Scales in Quantum Field Theory	231
12.2	Dimensional Regularization	232
12.3	The ϕ^4 Theory: A Case Study	236
12.4	The Renormalization Group Equations in Dimensional Regularization	244
12.5	The Issue of Quadratic Divergences	247
12.6	Effective Field Theories: A Brief Introduction	249
12.7	Remarks on Naturalness	255
12.8	Coda: Heavy Particles and Decoupling	257
	References	260
13	Special Topics	261
13.1	Creation of Particles by Classical Fields	261
13.2	Supersymmetry	268
	References	273
	Appendix A: Notation, Conventions and Units	275
	Appendix B: A Crash Course in Group Theory	277
	Index	289

Chapter 1

Why Do We Need Quantum Field Theory After All?

Quantum field theory is the basic tool to understand the physics of the elementary constituents of matter (see [1–15] for an incomplete list of textbooks in the subject). It is both a very powerful and a very precise framework: using it we can describe physical processes in a range of energies going from the few millions electrovolts typical of nuclear physics to the thousands of billions of the Large Hadron Collider (LHC). And all this with astonishing precision.

In this first chapter our aim is to explain why quantum mechanics is not enough and how quantum field theory is forced upon us by special relativity. We will review a number of riddles that appear in the attempt to extend the results of quantum mechanics to systems where relativistic effects cannot be ignored. Their resolution requires giving up the quantum mechanical description of a single particle to allow for the creation and annihilation of particles. As we will see, quantum fields provide the right tool to handle this.

1.1 Relativistic Quantum Mechanics

In spite of the impressive success of quantum mechanics in describing atomic physics, it was immediately clear after its formulation that its relativistic extension was not free of difficulties. These problems were clear already to Schrödinger, whose first guess for a wave equation of a free relativistic particle was the Klein–Gordon equation¹

$$\left(\frac{\partial^2}{\partial t^2} - \nabla^2 + m^2 \right) \psi(t, \mathbf{x}) = 0. \quad (1.1)$$

This equation follows directly from the relativistic “mass-shell” identity $E^2 = \mathbf{p}^2 + m^2$ using the correspondence principle

¹ We use natural units $\hbar = c = 1$. A summary of the units and conventions used in the book can be found in Appendix A.

$$\begin{aligned} E &\rightarrow i \frac{\partial}{\partial t}, \\ \mathbf{p} &\rightarrow -i \nabla. \end{aligned} \quad (1.2)$$

Plane wave solutions to the wave equation (1.1) are readily obtained

$$\psi(t, \mathbf{x}) = e^{-ip_\mu x^\mu} = e^{\mp i E_{\mathbf{p}} t + i \mathbf{p} \cdot \mathbf{x}} \quad (1.3)$$

with

$$E_{\mathbf{p}} = \sqrt{\mathbf{p}^2 + m^2}. \quad (1.4)$$

In order to have a complete basis of functions, we must include both signs in the exponent. The probability density is read from the time component of the conserved current

$$j_\mu = \frac{i}{2} (\psi^* \partial_\mu \psi - \partial_\mu \psi^* \psi), \quad (1.5)$$

Since $j^0 = E$, we find that it is not positive definite.

A complete, properly normalized, continuous basis of solutions of the Klein–Gordon equation (1.1) labelled by the momentum \mathbf{p} is given by

$$\begin{aligned} f_{\mathbf{p}}(t, \mathbf{x}) &= \frac{1}{(2\pi)^{\frac{3}{2}} \sqrt{2E_{\mathbf{p}}}} e^{-i E_{\mathbf{p}} t + i \mathbf{p} \cdot \mathbf{x}}, \\ f_{-\mathbf{p}}(t, \mathbf{x}) &= \frac{1}{(2\pi)^{\frac{3}{2}} \sqrt{2E_{\mathbf{p}}}} e^{i E_{\mathbf{p}} t - i \mathbf{p} \cdot \mathbf{x}}. \end{aligned} \quad (1.6)$$

Defining the inner product

$$\langle \psi_1 | \psi_2 \rangle = i \int d^3x (\psi_1^* \partial_0 \psi_2 - \partial_0 \psi_1^* \psi_2),$$

the states (1.6) form an orthonormal basis

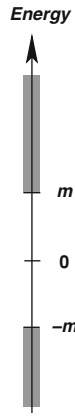
$$\begin{aligned} \langle f_{\mathbf{p}} | f_{\mathbf{p}'} \rangle &= \delta(\mathbf{p} - \mathbf{p}'), \\ \langle f_{-\mathbf{p}} | f_{-\mathbf{p}'} \rangle &= -\delta(\mathbf{p} - \mathbf{p}'), \end{aligned} \quad (1.7)$$

$$\langle f_{\mathbf{p}} | f_{-\mathbf{p}'} \rangle = 0. \quad (1.8)$$

The wave functions $f_{\mathbf{p}}(t, \mathbf{x})$ describe states with momentum \mathbf{p} and energy $E_{\mathbf{p}} = \sqrt{\mathbf{p}^2 + m^2}$. On the other hand, the wave functions $f_{-\mathbf{p}}(t, \mathbf{x})$ not only have negative scalar product but they correspond to negative energy states

$$i \frac{\partial}{\partial t} f_{-\mathbf{p}}(t, \mathbf{x}) = -\sqrt{\mathbf{p}^2 + m^2} f_{-\mathbf{p}}(t, \mathbf{x}). \quad (1.9)$$

Fig. 1.1 Spectrum of the Klein–Gordon wave equation



Therefore the energy spectrum of the theory satisfies $|E| > m$ and is unbounded from below (see Fig. 1.1). Although in the case of a free theory the absence of a ground state is not necessarily a fatal problem, once the theory is coupled to the electromagnetic field this is the source of all kinds of disasters, since nothing can prevent the decay of any state by the emission of electromagnetic radiation.

The problem of the instability of the “first-quantized” relativistic wave equation can be heuristically tackled in the case of spin- $\frac{1}{2}$ particles, described by the Dirac equation

$$\left(-i\beta\frac{\partial}{\partial t} + \alpha \cdot \nabla - m\right)\psi(t, \mathbf{x}) = 0, \quad (1.10)$$

where α and β are 4×4 matrices

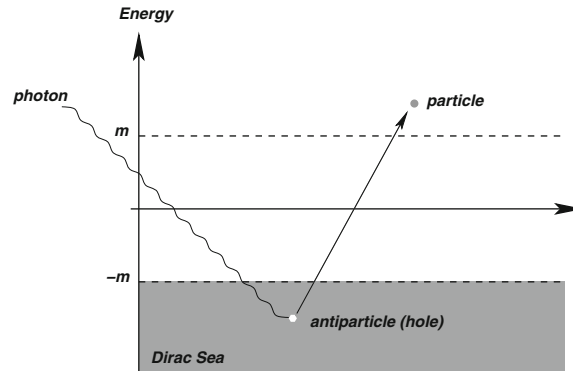
$$\alpha^i = \begin{pmatrix} 0 & i\sigma_i \\ -i\sigma_i & 0 \end{pmatrix}, \quad \beta = \begin{pmatrix} 0 & \mathbf{1} \\ \mathbf{1} & 0 \end{pmatrix}, \quad (1.11)$$

with σ_i the Pauli matrices (see Appendix A) and the wave function $\psi(t, \mathbf{x})$ has four components: it is a Dirac spinor, an object that will be studied in more detail in Chap. 3. The wave equation (1.10) can be thought of as a kind of “square root” of the Klein–Gordon equation (1.1), since the latter can be obtained as

$$\begin{aligned} & \left(-i\beta\frac{\partial}{\partial t} + \alpha \cdot \nabla - m\right)^\dagger \left(-i\beta\frac{\partial}{\partial t} + \alpha \cdot \nabla - m\right)\psi(t, \mathbf{x}) \\ &= \left(\frac{\partial^2}{\partial t^2} - \nabla^2 + m^2\right)\psi(t, \mathbf{x}). \end{aligned} \quad (1.12)$$

An analysis of Eq. (1.10) along the lines of the one presented for the Klein–Gordon equation leads again to the existence of negative energy states and

Fig. 1.2 Creation of a particle-antiparticle pair in the Dirac sea picture



a spectrum unbounded from below as in Fig. 1.1. Dirac, however, solved the instability problem by pointing out that now the particles are fermions and therefore they are subject to Pauli’s exclusion principle. Hence, each state in the spectrum can be occupied by at most one particle, so the states with $E = m$ can be made stable if we assume that *all* the negative energy states are filled.

If Dirac’s idea restores the stability of the spectrum by introducing a stable vacuum where all negative energy states are occupied, the so-called Dirac sea, it also leads directly to the conclusion that a single-particle interpretation of the Dirac equation is not possible. Indeed, a photon with enough energy ($E > 2m$) can excite one of the electrons filling the negative energy states, leaving behind a “hole” in the Dirac sea (see Fig. 1.2). This hole behaves as a particle with equal mass and opposite charge that is interpreted as a positron, so there is no escape to the conclusion that interactions will produce particle-antiparticle pairs out of the vacuum.

1.2 The Klein Paradox

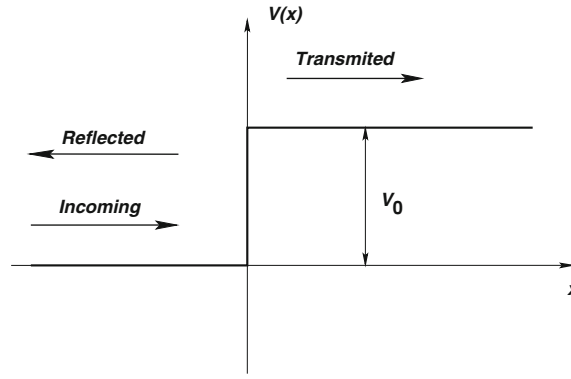
In spite of the success of the heuristic interpretation of negative energy states in the Dirac equation this is not the end of the story. In 1929 Oskar Klein stumbled into an apparent paradox when trying to describe the scattering of a relativistic electron by a square potential using Dirac’s wave equation [16] (for pedagogical reviews see [17–19]). In order to capture the essence of the problem without entering into unnecessary complication we will study Klein’s paradox in the context of the Klein–Gordon equation.

Let us consider a square potential with height $V_0 > 0$ of the type showed in Fig. 1.3. A solution to the wave equation in regions I and II is given by

$$\begin{aligned}\psi_I(t, x) &= e^{-iEt+ip_1x} + R e^{-iEt-ip_1x}, \\ \psi_{II}(t, x) &= T e^{-iEt+ip_2x},\end{aligned}\tag{1.13}$$

where the mass-shell condition implies

Fig. 1.3 Illustration of the Klein paradox



$$p_1 = \sqrt{E^2 - m^2}, \quad p_2 = \sqrt{(E - V_0)^2 - m^2}. \quad (1.14)$$

The constants R and T are computed by matching the two solutions across the boundary $x = 0$. The conditions $\psi_I(t, 0) = \psi_{II}(t, 0)$ and $\partial_x \psi_I(t, 0) = \partial_x \psi_{II}(t, 0)$ imply that

$$T = \frac{2p_1}{p_1 + p_2}, \quad R = \frac{p_1 - p_2}{p_1 + p_2}. \quad (1.15)$$

At first sight one would expect a behavior similar to the one encountered in the nonrelativistic case. If the kinetic energy is bigger than V_0 both a transmitted and reflected wave are expected, whereas when the kinetic energy is smaller than V_0 one only expects to find a reflected wave, the transmitted wave being exponentially damped within a distance of a Compton wavelength inside the barrier.

This is indeed what happens if $E - m > V_0$. In this case both p_1 and p_2 are real and we have a partly reflected, and a partly transmitted wave. In the same way, if $V_0 - 2m < E - m < V_0$ then p_2 is imaginary and there is total reflection.

However, in the case when $V_0 > 2m$ and the energy is in the range $0 < E - m < V_0 - 2m$ a completely different situation arises. In this case one finds that both p_1 and p_2 are real and therefore the incoming wave function is partially reflected and partially transmitted across the barrier. This is a shocking result, since it implies that there is a nonvanishing probability of finding the particle at any point across the barrier with negative kinetic energy ($E - m - V_0 < 0$)! This weird result is known as Klein's paradox.

As with the negative energy states, the Klein paradox results from our insistence in giving a single-particle interpretation to the relativistic wave function. In fact, a multiparticle analysis of the paradox [17] shows that what happens when $0 < E - m < V_0 - 2m$ is that the reflection of the incoming particle by the barrier is accompanied by the creation of particle-antiparticle pairs out of the energy of the barrier (notice that the condition implies that $V_0 > 2m$, the threshold for the creation of a particle-antiparticle pair).

This particle creation can be understood by noticing that the sudden potential step in Fig. 1.3 localizes the incoming particle with mass m in distances smaller than its Compton wavelength $\lambda = 1/m$. This can be seen by replacing the square potential by another one where the potential varies smoothly from 0 to $V_0 > 2m$ in distance scales larger than $1/m$. This case was worked out by Sauter shortly after Klein pointed out the paradox [20]. He considered a situation where the regions with $V = 0$ and $V = V_0$ are connected by a region of length d with a linear potential $V(x) = V_0 x/d$. When $d > 1/m$ he found that the transmission coefficient is exponentially small.²

1.3 From Wave Functions to Quantum Fields

The creation of particles is impossible to avoid whenever one tries to localize a particle of mass m within its Compton wavelength. Indeed, from the Heisenberg uncertainty relation we find that if $\Delta x \sim 1/m$, the fluctuations in the momentum will be of order $\Delta p \sim m$ and fluctuations in the energy of order

$$\Delta E \sim m \quad (1.16)$$

can be expected. Therefore, in a relativistic theory, the fluctuations of the energy are enough to allow for the creation of particles out of the vacuum. In the case of a spin- $\frac{1}{2}$ particle, the Dirac sea picture shows clearly how, when the energy fluctuations are of order m , electrons from the Dirac sea can be excited to positive energy states, thus creating electron–positron pairs.

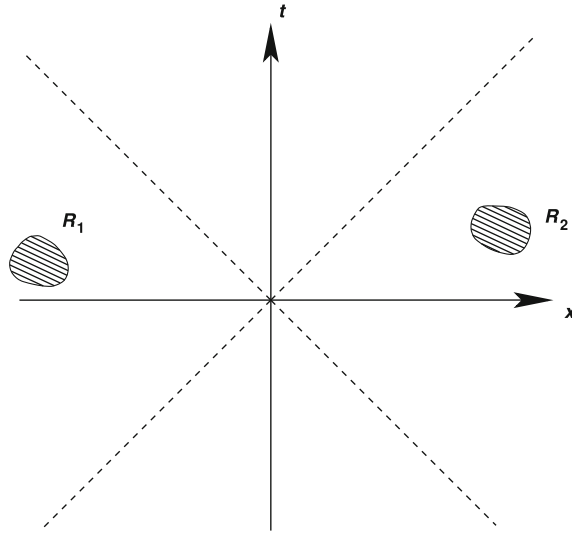
It is possible to see how the multiparticle interpretation is forced upon us by relativistic invariance. In non-relativistic quantum mechanics observables are represented by self-adjoint operator that in the Heisenberg picture depend on time. Therefore measurements are localized in time but are global in space. The situation is radically different in the relativistic case. Since no signal can propagate faster than the speed of light, measurements have to be localized both in time and space. Causality demands then that two measurements carried out in causally-disconnected regions of space–time cannot interfere with each other. In mathematical terms this means that if \mathcal{O}_{R_1} and \mathcal{O}_{R_2} are the observables associated with two measurements localized in two causally-disconnected regions R_1, R_2 (see Fig. 1.4), they satisfy

$$[\mathcal{O}_{R_1}, \mathcal{O}_{R_2}] = 0, \quad \text{if } (x_1 - x_2)^2 < 0, \quad \text{for all } x_1 \in R_1, \quad x_2 \in R_2. \quad (1.17)$$

Hence, in a relativistic theory, the basic operators in the Heisenberg picture must depend on the space–time position x^μ . Unlike the case in non-relativistic quantum mechanics, here the position \mathbf{x} is *not* an observable, but just a label, similarly to the case of time in ordinary quantum mechanics. Causality is then imposed microscopically by requiring

² In Sect. 13.1 we will see how, in the case of the Dirac field, this exponential behavior can be associated with the creation of electron–positron pairs due to a constant electric field (Schwinger effect).

Fig. 1.4 Two regions R_1 , R_2 that are causally disconnected



$$[\mathcal{O}(x), \mathcal{O}(y)] = 0, \quad \text{if } (x - y)^2 < 0. \quad (1.18)$$

A smeared operator \mathcal{O}_R over a space-time region R can then be defined as

$$\mathcal{O}_R = \int d^4x \mathcal{O}(x) f_R(x) \quad (1.19)$$

where $f_R(x)$ is the characteristic function associated with R ,

$$f_R(x) = \begin{cases} 1 & x \in R \\ 0 & x \notin R \end{cases}. \quad (1.20)$$

Equation (1.17) follows now from the microcausality condition (1.18).

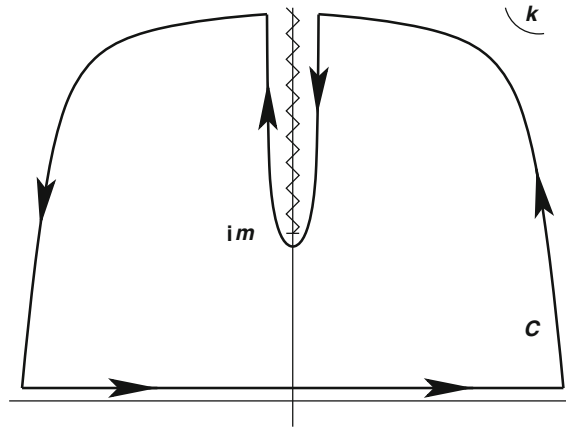
Therefore, relativistic invariance forces the introduction of quantum fields. It is only when we insist in keeping a single-particle interpretation that we crash against causality violations. To illustrate the point, let us consider a single particle wave function $\psi(t, \mathbf{x})$ that initially is localized in the position $\mathbf{x} = 0$

$$\psi(0, \mathbf{x}) = \delta(\mathbf{x}). \quad (1.21)$$

Evolving this wave function using the Hamiltonian $H = \sqrt{-\nabla^2 + m^2}$ we find that the wave function can be written as

$$\psi(t, \mathbf{x}) = e^{-it\sqrt{-\nabla^2 + m^2}} \delta(\mathbf{x}) = \int \frac{d^3k}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{x} - it\sqrt{k^2 + m^2}}. \quad (1.22)$$

Fig. 1.5 Complex contour C for the computation of the integral in Eq. (1.23)



Integrating over the angular variables, the wave function can be recast in the form

$$\psi(t, \mathbf{x}) = \frac{-i}{4\pi^2|\mathbf{x}|} \int_{-\infty}^{\infty} k dk e^{ik|\mathbf{x}|} e^{-it\sqrt{k^2+m^2}}. \quad (1.23)$$

The resulting integral can be evaluated using the complex integration contour C shown in Fig. 1.5. The result is that, for any $t > 0$, $\psi(t, \mathbf{x}) \neq 0$ for any \mathbf{x} . If we insist in interpreting the wave function $\psi(t, \mathbf{x})$ as the probability density of finding the particle at the location \mathbf{x} at the time t , the probability leaks out of the light cone, thus violating causality.

The bottom line of the analysis of this chapter is clear: a fully relativistic quantum theory must give up the idea of describing the system in terms of the wave function of a *single* particle. As a matter of fact, relativistic quantum mechanics is, at best, a narrow boundary area. It might be a useful tool to compute the first relativistic corrections in certain quantum systems. However it runs into serious trouble as soon as one tries to use it for a full-fledged relativistic description of the quantum phenomena. Next we will see how quantum field theory provides the right framework to handle these problems.

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