#### **Quantisation**

between electron and positron creation-annihilation operators We repeat the bosonic arguments, except for the fact that we have now anti-commutation relations

$$\widehat{\psi}_{\scriptscriptstylearphi}(t,ec{x}) = \sum_{\scriptscriptstylearphi=\pmrac{1}{2}} \int rac{d^3k}{(2\pi)^3} rac{1}{2\omega_{ec{k}}} \left[ u_{\scriptscriptstylearphi}(ec{k},s) \widehat{b}(ec{k},s) e^{-i\omega_{ec{k}}t+iec{k}\cdotec{x}} + v_{\scriptscriptstylearphi}(ec{k},s) \widehat{d}^{\,\dagger}(ec{k},s) e^{i\omega_{ec{k}}t-iec{k}\cdotec{x}} 
ight].$$

$$\{\hat{\psi}_{\alpha}(t,\mathbf{x}),\hat{\psi}_{\beta}^{\dagger}(t,\mathbf{y})\} = \delta(\mathbf{x} - \mathbf{y})\delta_{\alpha\beta}$$

$$\{b(\mathbf{k},s),b^{\dagger}(\mathbf{k'},s')\}=(2\pi)^{3}(2\omega_{\mathbf{k}})\delta(\mathbf{k-k'})\delta_{ss'}$$

$$\{b(\mathbf{k},s),b(\mathbf{k'},s')\} = \{b^{\dagger}(\mathbf{k},s),b^{\dagger}(\mathbf{k'},s')\} = 0.$$
$$\{d(\mathbf{k},s),d^{\dagger}(\mathbf{k'},s')\} = (2\pi)^{3}(2\omega_{\mathbf{k}})\delta(\mathbf{k}-\mathbf{k'})\delta_{ss'}.$$

$$\widehat{H} = \frac{1}{2} \sum_{s=\pm \frac{1}{2}} \int \frac{d^3k}{(2\pi)^3} \left[ b^{\dagger}(\mathbf{k}, s) b(\mathbf{k}, s) - d(\mathbf{k}, s) d^{\dagger}(\mathbf{k}, s) \right].$$

$$\{d(\mathbf{k},s),d(\mathbf{k'},s')\} = \{d^{\dagger}(\mathbf{k},s),d^{\dagger}(\mathbf{k'},s')\} = 0.$$

$$\widehat{H} = \sum_{ec{s}=\pm\pm} \int rac{d^3k}{(2\pi)^3} rac{1}{2\omega_{ec{k}}} \left[ \omega_{ec{k}} b^\dagger(ec{k},s) b(ec{k},s) + \omega_{ec{k}} d^\dagger(ec{k},s) d(ec{k},s) 
ight] - \ 2 \int d^3k \, \omega_{ec{k}} \delta(ec{0}).$$

We have a conserved charge and current

$$j^{\mu} = \overline{\psi}\gamma^{\mu}\psi,$$

$$\hat{c}_{\mu}j^{\mu}=0$$

$$Q = e \int d^3x j^0$$

The two-point function or Feynman propagator is:

$$S_{\alpha\beta}(x_1, x_2) = \langle 0|T \left[ \psi_{\alpha}(x_1) \overline{\psi}_{\beta}(x_2) \right] |0\rangle$$

$$T \left[ \psi_{\alpha}(x) \overline{\psi}_{\beta}(y) \right] = \theta(x^0 - y^0) \psi_{\alpha}(x) \overline{\psi}_{\beta}(y) - \theta(y^0 - x^0) \overline{\psi}_{\beta}(y) \psi_{\alpha}(x).$$



### Introducing gauge fields

invariant. This is not so in the quantum theory, where we need to use the vector and scalar theory can be formulated exclusively in terms of the E,B field that are manifestly gauge a gauge field. For some time this symmetry sounded like a luxury. In fact the classical Aharonov effect and the quantisation of electric charge (if there is a single monopole in the The canonical gauge field is the electromagnetic field. The first one that was understood as Universe, (Dirac)). potentials. There are new, non-local observables. They are responsible for the Bohm-

suitable generalisations of the EM field. They are gauge theories. In fact it seems as though low-energy accidents. All symmetries in the UV should be local. What we have learned is that all fundamental interactions known to us are mediated by Nature abhors global symmetries. It appears that all the known global symmetries are just

finds an explanation. Unfortunately there is no evidence for it at this moment... We do not know why this should be so. String Theory is the only theory where this fact

## E&M in Quantum Mechanics

#### Classical EM

$$\nabla \cdot \mathbf{E} = 0$$

$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial}{\partial t} \mathbf{B}$$

$$\nabla \times \mathbf{B} = \frac{\partial}{\partial t} \mathbf{E}$$

$$= -\nabla \varphi - \frac{\partial \mathbf{A}}{\partial t}$$
$$= \nabla \times \mathbf{A}.$$

$$\begin{split} \partial_{\mu}F^{\mu\nu} &= j^{\mu} \quad j^{\mu} = (\rho_{\cdot}\mathbf{j}) \\ \varepsilon^{\mu\nu\sigma\eta}\partial_{\nu}F_{\sigma\eta} &= 0, \quad A^{\mu} = (\varphi_{\cdot}\mathbf{A}) \\ F_{\mu\nu} &= \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} \end{split}$$

Classical EM in relativistic form

function, this gives subtle consequences to gauge symmetry Coupling to QM requires the gauge potentials and a non-trivial transformation of the wave

$$i\frac{\partial}{\partial t}\Psi = \left[ -\frac{1}{2m} (\nabla - ie\mathbf{A})^2 + e\varphi \right] \Psi$$

$$\Psi(t, \mathbf{x}) \longrightarrow e^{-ie\varepsilon(t, \mathbf{x})} \Psi(t, \mathbf{x})$$

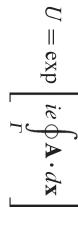
$$\varphi(t, \mathbf{x}) \rightarrow \varphi(t, \mathbf{x}) + \frac{\partial}{\partial t} \varepsilon(t, \mathbf{x}), \quad \mathbf{A}(t, \mathbf{x}) \rightarrow \mathbf{A}(t, \mathbf{x}) + \nabla \varepsilon(t, \mathbf{x}).$$

$$A_{\mu} \longrightarrow A_{\mu} + \partial_{\mu} \varepsilon$$



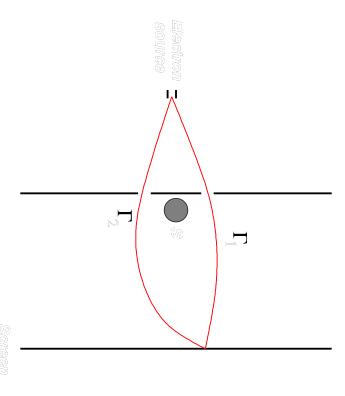
### Non-local observables

$$\Psi = e^{ie \int_{\Gamma_1} \mathbf{A} \cdot d\mathbf{x}} \Psi_1^{(0)} + e^{ie \int_{\Gamma_2} \mathbf{A} \cdot d\mathbf{x}} \Psi_2^{(0)}$$
$$= e^{ie \int_{\Gamma_1} \mathbf{A} \cdot d\mathbf{x}} \left[ \Psi_1^{(0)} + e^{ie \oint_{\Gamma} \mathbf{A} \cdot d\mathbf{x}} \Psi_2^{(0)} \right]$$



This is the Aharonov-Bohm effect.

The phase factor, and its non-abelian



involves E,B hence the electrons would not see the solenoid at all!! generalisation are known as "Wilson loops" or holonomies of the gauge field. Note that classically there would be no effect. The Lorentz force equation only  $m\frac{du^{\mu}}{d\tau} = eF^{\mu\nu}u_{\nu}$ 

$$n\frac{du^{r}}{d\tau} = eF^{\mu\nu}u_{\nu}$$



#### Dirac and charge quantisation Magnetic monopoles:

$$7 \cdot \mathbf{E} = 0$$

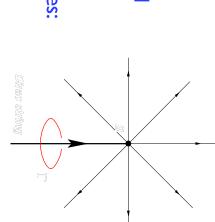
$$\nabla \cdot \mathbf{B} = 0$$

$$\nabla \times \mathbf{E} = -\frac{\partial}{\partial t} \mathbf{B}$$

$$\nabla \times \mathbf{B} = \frac{\partial}{\partial t} \mathbf{E}$$

$$\mathbf{E} - i\mathbf{B} \longrightarrow e^{i\theta} (\mathbf{E} - i\mathbf{B})$$

For angle = 90 E and B get exchanged



The symmetry extend to matter if we have magnetic sources:

$$\rho - i \rho_m \longrightarrow e^{i\theta} (\rho - i \rho_m), \qquad \mathbf{j} - i \mathbf{j}_m \longrightarrow e^{i\theta} (\mathbf{j} - i \mathbf{j}_m).$$

$$-i\mathbf{j}_m \longrightarrow e^{i\theta}(\mathbf{j}-i\mathbf{j}_m)$$

Consider a magnetic pole:

$$\nabla \cdot \mathbf{B} = g \, \delta(\mathbf{x}).$$

$$B_r = \frac{1}{4\pi} \frac{g}{|\mathbf{x}|^2},$$

$$B_{\varphi} = B_{\theta} = 0$$

$$A_{\varphi} = \frac{1}{4\pi} \frac{g}{|\mathbf{x}|} \tan \frac{\theta}{2},$$

$$A_r = A_\theta = 0.$$

unobservable. Then we can do a "A-B" like argument (Dirac did it 20 years earlier). We should not forget the fact that there is a factor of The Dirac string can be changed by gauge transformations, in doing QM it has to be  $e^{ieg}=1$  $eg = 2\pi n$ 

$$e^{ieg} = 1$$

$$q_1g_2 - q_2g_1 = 2\pi n,$$



# Electromagnetic Fields and Photons

Ignoring sources, the E&M field is a "free field"

ring sources, the E&M field is a "free field" 
$$\mathbf{E} = -\nabla \varphi - \frac{\partial \mathbf{A}}{\partial t}$$
 
$$\mathcal{L}_{\text{Maxwell}} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} = \frac{1}{2} \left( \mathbf{E}^2 - \mathbf{B}^2 \right).$$
 
$$\mathbf{B} = \nabla \times \mathbf{A}.$$

$$F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu} \qquad A_{\mu} \longrightarrow A_{\mu} + \partial_{\mu}\varepsilon$$

"coordinate" q

vector potential the momentum p and the

The electric field is the

$$0 = \partial_{\mu}\partial^{\mu}A^{\nu} - \partial_{\nu}\left(\partial_{\mu}A^{\mu}\right) = \partial_{\mu}\partial^{\mu}A^{\nu}$$

To be able to invert, we need to fix the gauge:

$$\partial_{\mu}A^{\mu}=0.$$

As usual, we look for plane wave solutions fix the gauge Residual gauge transformation used to fully

$$\varepsilon_{\mu}(\mathbf{k},\lambda)e^{-i|\mathbf{k}|t+i\mathbf{k}\cdot\mathbf{x}}$$

 $k^{\mu} \varepsilon_{\mu}(\mathbf{k}, \lambda) = 0$ 

$$k^2 = k_{\mu}k^{\mu} = (k^0)^2 - \mathbf{k}^2 = 0$$

 $\epsilon_{\mu}(\mathbf{k},\lambda) \to \epsilon_{\mu}(\mathbf{k},\lambda) + k_{\mu} \chi(\mathbf{k}), \ k^2 = 0$ 

polarisations. Gauge invariance seems more a redundancy rather than a symmetry in the description of the theory Now, as usual we expand the field in oscillator and apply CCR. After fully fixing the gauge there are only two physical

$$\widehat{A}_{\mu}(t,\mathbf{x}) = \sum_{\lambda = \pm 1} \int \frac{d^3k}{(2\pi)^3} \frac{1}{2|\mathbf{k}|} \left[ \varepsilon_{\mu}(\mathbf{k},\lambda) \widehat{a}(\mathbf{k},\lambda) e^{-i|\mathbf{k}|t + i\mathbf{k} \cdot \mathbf{x}} + \varepsilon_{\mu}(\mathbf{k},\lambda)^* \widehat{a}^{\dagger}(\mathbf{k},\lambda) e^{i|\mathbf{k}|t - i\mathbf{k} \cdot \mathbf{x}} \right].$$

$$\left[\widehat{a}(\mathbf{k},\lambda),\widehat{a}^{\dagger}(\mathbf{k'},\lambda')\right] = (2\pi)^{3}(2|\mathbf{k}|)\delta(\mathbf{k}-\mathbf{k'})\delta_{\lambda\lambda'}$$

then we get negative probabilities (Gupta-Bleuler, BRST) If we keep all four polarisation by partial gauge fixing,

$$\delta_{\mathbb{A},\mathbb{A}'} o -\eta_{\mathbb{A},\mathbb{A}'}$$



#### Coupling matter

coupling. We make derivatives covariant with respect to space-time dependent changes of phases We imitate the coupling in the Schrödinger equation, this is what used to be called minimal in the wave-function

$$i\frac{\partial}{\partial t}\Psi = \left[ -\frac{1}{2m} (\nabla - ie\mathbf{A})^2 + e\varphi \right] \Psi \qquad D_{\mu} \left[ e^{ie\varepsilon(x)} \psi \right] = e^{ie\varepsilon(x)} D_{\mu} \psi.$$

$$\Psi(t, \mathbf{x}) \longrightarrow e^{-ie\varepsilon(t, \mathbf{x})} \Psi(t, \mathbf{x}) \qquad D_{\mu} = \partial_{\mu} - ieA_{\mu}.$$

$$A_{\mu} \longrightarrow A_{\mu} + \partial_{\mu}\varepsilon$$

phase rotations, a physically more satisfactory concept. This defines the coupling of the electron to the E&M field The rigid phase rotation invariance of the Dirac Lagrangian for electrons is transformed into local

$$\mathscr{L}_{\mathrm{QED}} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \overline{\psi} (i \not \! D - m) \psi, \qquad \mathscr{L}_{\mathrm{QED}}^{\mathrm{(int)}} = -e A_{\mu} \, \overline{\psi} \gamma^{\mu} \psi.$$

$$\psi \longrightarrow e^{ie\varepsilon(x)}\psi, \qquad A_{\mu} \longrightarrow A_{\mu} + \partial_{\mu}\varepsilon(x).$$

electron. This is QED, the best tested theory in the history of science, an example is the gyromagnetic ratio of the

$$grac{e}{8m}[\gamma^{\&},\gamma^{\swarrow}]F_{\mu
u}$$

$$g/2 = 1.00115965218085(76)$$
  $\vec{\mu} = g_{\mu} \frac{e\hbar}{2m_{\mu}c} \vec{s},$   $\alpha^{-1} = 137.035999070(98)$ 

$$=g_{\mu}\frac{en}{2m_{\mu}c}\vec{s}, \qquad g_{\mu}=2(1+a_{\mu})$$
Dirac

## Group Theory reminder

For the SM all group we will need are:

$$[T^{\scriptscriptstyle ilde{arphi}},T^{\scriptscriptstyle ilde{arphi}}]=if^{\scriptscriptstyle ilde{arphi}\circ}T^{\scriptscriptstyle ilde{arphi}}$$

$$G_{SM} = SU(3) \times SU(2) \times U(1)$$

$$g \in G$$

$$g=e^{iarepsilon^{lpha}}$$
 The

$$\operatorname{tr}(T^{\scriptscriptstyle ilde{m{arphi}}}) = T_{\scriptscriptstyle ilde{m{arphi}}}(R) \, \delta^{\scriptscriptstyle ilde{m{arphi}}}$$

$$\det g = 1 \Rightarrow \operatorname{tr} T^a = 0$$
 (for SU(2), SU(3) not for U(1) of course)

SU(2): angular momentum, isospin, and also weak isospin

 $\mathsf{U}(\mathsf{I})$  is of course the simplest, just phase multiplication, i.e. as in QED

$$[T^a, T^b] = i\varepsilon^{abc}T^c$$
,  $T^{\pm} = \frac{1}{\sqrt{2}}(T^1 \pm iT^2)$ ,  $T^3$ 

$$=\frac{1}{\sqrt{2}}(T^1\pm iT^2),$$

$$[T^3, T^{\pm}] = \pm T^{\pm}$$
  
 $[T^+, T^-] = T^3$ 

$$T^a=rac{1}{2}\sigma^a$$
 For spin ½  ${
m tr}rac{\sigma^{\scriptscriptstyle ilde lpha}}{2}rac{1}{2}=rac{1}{2}\delta^{\scriptscriptstyle ilde a \delta}$ 

$$a,b=1,2,3$$

$$J^{1} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & -1 & 0 \end{pmatrix}, \quad J^{2} = \begin{pmatrix} 0 & 0 & -1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}, \quad J^{3} = \begin{pmatrix} 0 & 1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

For SU(3) the generators are the eight Gell-Mann 3x3 traceless hermitean matrices chosen to satisfy:

$$\mathrm{tr}rac{\lambda^a}{2}rac{\lambda^a}{2} \,=\, rac{1}{2}\delta^{ab};\; a,b=1,\ldots,8$$

SU(3) of color, an exact gauge symmetry, also flavor SU(3), which is global (see later)



#### More about SU(3)

for color SU(3): There are very few representations we will need

quarks

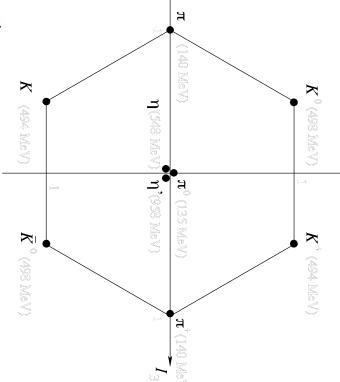
antiquarks gluons

For flavor SU(3) more needed: mesons,

baryons

 $3, \overline{3}, 8, 10, 10, 27 \dots$ 

it was theorised earlier. an approximate global symmetry of the problem, although compatible with color gauge symmetry, flavor appears as A remarkable fact about the SM and QCD in particular is the fact that once we write the most general Lagrangian



pseudo-scalar meson octet

$$Q = I_3 + \frac{B+S}{2},$$

$$|\Delta^{++}; s_z = \frac{3}{2}\rangle = |uuu\rangle \otimes |\uparrow\uparrow\uparrow\rangle \equiv |u\uparrow, u\uparrow, u\uparrow\rangle$$

$$|uud\rangle_S = \frac{1}{\sqrt{6}} \Big( |uud\rangle + |udu\rangle - 2|duu\rangle \Big), \qquad |\uparrow\uparrow\rangle_S = \frac{1}{\sqrt{6}} \Big( |\uparrow\uparrow\downarrow\rangle + |\uparrow\downarrow\uparrow\rangle - 2|\uparrow\uparrow\downarrow\rangle \Big).$$

 $|uud\rangle_A = \frac{1}{\sqrt{2}} (|uud\rangle - |udu\rangle).$ 

$$\rangle), \qquad |\uparrow\uparrow\rangle_S = \frac{1}{\sqrt{6}} (|\uparrow\uparrow\downarrow\rangle + |\uparrow\downarrow\uparrow\rangle - 2|\uparrow\uparrow\downarrow$$

 $|\uparrow\uparrow\rangle_A = \frac{1}{\sqrt{2}} (|\uparrow\downarrow\uparrow\rangle - |\downarrow\uparrow\uparrow\rangle).$ 

$$|p\downarrow\rangle = \frac{1}{\sqrt{2}} \left( |uud\rangle_S \otimes |\downarrow\downarrow\rangle_A + |uud\rangle_A \otimes |\downarrow\downarrow\rangle_S \right).$$

 $|p\uparrow\rangle = \frac{1}{\sqrt{2}} (|uud\rangle_S \otimes |\uparrow\uparrow\rangle_A + |uud\rangle_A \otimes |\uparrow\uparrow\rangle_S).$ 



# Gauge theories and their quantisation

Imagine we have a theory with a global symmetry

$$\psi o g \, \psi$$

$$ar{\psi} 
ightarrow ar{\psi} g^\dagger$$

$${\cal L}=ar{\psi}ioldsymbol{\phi}\,\psi$$

Imitating electromagnetism:

$$\partial_{\mu} \to D_{\mu} \psi = (\partial_{\mu} + ieA_{\mu}^{a} T^{a}) \psi \equiv (\partial_{\mu} + ieA_{\mu}) \psi$$

$$D_{\mu}\psi o gD_{\mu}\psi$$

We can read off the gauge field transformations

$$A_{\mu} \to \frac{1}{ie} g \,\partial_{\mu} g^{-1} + g A_{\mu} g^{-1}$$

$$g \approx 1 + \epsilon \qquad A_{\mu} \to A_{\mu} + \frac{1}{ie} D_{\mu} \epsilon \qquad D_{\mu} \epsilon + ie[A_{\mu}, \epsilon]$$

$$[D_{\mu}, D_{\nu}] = ieT^a F^a_{\mu\nu},$$

$$egin{aligned} &+ rac{\dot{ar{i}}}{ie} D_{\mu} \epsilon & D_{\mu} \epsilon + ie[A_{\mu}, \epsilon] \ &F^a_{\mu 
u} = \partial_{\mu} A^a_{
u} - \partial_{
u} A^a_{\mu} - e f^{abc} A^b_{\mu} A^c_{
u} \end{aligned}$$

$$F_{\mu\nu} \equiv T^a F^a_{\mu\nu} \to g F_{\mu\nu} g^{-1}$$

responsible for confinement, among other things Nonabelian gauge fields have self-couplings unlike photons. This is

General gauge theory Lagrangian:

$$\mathcal{L} = -\frac{1}{4} F^a_{\mu\nu} F^{\mu\nu a} + i \overline{\psi} \not D \psi + (D_{\mu} \phi)^{\dagger} D^{\mu} \phi$$

$$-\overline{\psi}\left[M_1(\phi)+i\gamma_5M_2(\phi)\right]\psi-V(\phi).$$

We need to provide the gauge group and the matter representations for bosons and fermions and off we go

Quantising a gauge theory is no joke. There are plenty of subtleties. We give you just a taste

We can define chromoelectric and magnetic fields as in QED

$$F_{0i}^{a} = \partial_{0}A_{i}^{a} - \partial_{i}A_{0}^{a} - ief^{abc}A_{0}^{b}A_{i}^{c} \equiv E_{i}^{a}$$
  
 $F_{ij}^{a} = \epsilon_{ijk}B_{k}^{a}, F_{0i}^{a} = \partial_{0}A_{0}^{a} - D_{i}A_{0}^{a}$ 

The canonical variables are 
$$oldsymbol{\Lambda}^{a} \mathbf{F}^{a}$$

$$\mathcal{L} = \mathbf{E}^a \, \partial_0 \mathbf{A}^a - rac{1}{2} (\mathbf{E}^2 + \mathbf{B}^2) - A_0^a \, (\mathbf{D} \cdot \mathbf{E})^a$$

 $A_0^a$  implements a constraint

We can read off the Hamiltonian density

### General Gauge Theory

$$H = \int d^3x \left(\frac{1}{2} (\mathbf{E}^2 + \mathbf{B}^2) + A_0^a (\mathbf{D} \cdot \mathbf{E})^a\right)$$

$$[A_i^a(\mathbf{x},0), E_j^b(\mathbf{y},0)] = i \, \delta_{ij} \, \delta^{ab} \delta(\mathbf{x} - \mathbf{y})$$

implemented as a constraint. motion, Gauss' law that has to be are missing one of the equations of in the Hamiltonian theory, but we independent gauge transformations that we only have time-We can fix the gauge A\_0=0 so

$$(\mathbf{D} \cdot \mathbf{E})^a = 0$$
 gauge trans

gauge transformations Cannot be implemented at the operator level. It generates

$$[Q(\epsilon), A_i^a] = i(D\epsilon)^a \qquad U(\epsilon) = \exp($$

$$[Q(\epsilon), A_i^a] = i(D\epsilon)^a$$
  $U(\epsilon) = \exp(i \int d^3x \, \epsilon^a(\mathbf{x}) \, (\mathbf{D} \cdot \mathbf{E})^a),$ 

 $U H U^{-1} = H$ 

Gauss' law becomes a condition on the physical states:

$$U(\epsilon)| ext{phys}
angle=| ext{phys}
angle$$

need choose only one element, this is done by Each gauge configuration sits in an orbit and we "fixing" the gauge for the t-independent gauge transt.

$$\mathbf{D} \cdot \mathbf{E} | \mathrm{phys} \rangle = 0$$





#### Some remarks

- Gauge symmetry is more a redundant description of the d.o.f.
- to the identity. Consider finite-E configurations Gauss' law implements gauge invariance under gauge t. connected

$$g(\mathbf{x}) = e^{i\alpha(\mathbf{x})} \to 1 \qquad |\mathbf{x}| \to \infty$$
 $\alpha(\mathbf{x}) \to 0 \qquad |\mathbf{x}| \to \infty$ 

There are others, and Gauss' law cannot impose invariance

$$g(\mathbf{x}): S^3 \to G, \ g(\infty) = 1$$

$$\pi_3(G)=Z$$
 the integers

$$g: S^1 \longrightarrow U(1),$$
  $g(x) = e^{i\alpha(x)}$   $\alpha(2\pi) = \alpha(0) + 2\pi n$ 



$$\oint_{S^1} g(x)^{-1} dg(x) = 2\pi n$$

$$n = \frac{1}{24\pi^2} \int_{S^3} d^3x \, \varepsilon_{ijk} \operatorname{Tr} \left[ \left( g^{-1} \partial_i g \right) \left( g^{-1} \partial_i g \right) \left( g^{-1} \partial_i g \right) \right]$$

You cannot comb a sphere



### A surprise: CP violation

- generated. This is a vacuum angle! In fact it violates CP. ❖ Gauge invariance only requires that under non-trivial transformations, a phase is
- It can be measured by looking for an edm of the neutron. So far no result:
- \* The strong CP problem, axions, invisible axions, axion cosmology, dark matter...

 $g_1 \in \mathcal{G}/\mathcal{G}_0$  the generator

$$\mathscr{U}(g_1)|\text{phys}\rangle = e^{i\theta}|\text{phys}\rangle.$$

$$S = -\frac{1}{4} \int d^4x F^a_{\mu\nu} F^{\mu\nu a} - \frac{\theta g_{\text{YM}}^2}{32\pi^2} \int d^4x F^a_{\mu\nu} \widetilde{F}^{\mu\nu a}$$

$$\widetilde{F}^a_{\mu\nu} = \frac{1}{2} \varepsilon_{\mu\nu\sigma\lambda} F^{\sigma\lambda a} \qquad F^a_{\mu\nu} \widetilde{F}^{\mu\nu a} = 4 \, \mathbf{E}^a \cdot \mathbf{B}^a$$

 $\frac{g_{\rm YM}^2}{32\pi^2} \int d^4x F_{\mu\nu}^a \widetilde{F}^{\mu\nu a}$ 

#### Computational tools

- the Path Integral formulation. There are computational rules in QFT: The canonical formalism and two general procedures to obtain
- integrations... Representation, Wick's theorem, You may recall that one used the Interaction T-products, Gaussian
- reminder a given scattering process, out of which we can evaluate the allow us to compute the probability amplitude associates to collider experiments. The next few pages provide simply a other quantities that can be observed for instance in decay width, differential and total cross section and many In the end we get a collection of well-defined rules that

### $\sqrt{p-m+i\varepsilon}$ $\int_{\beta\alpha}$

$$\mu \sim \nu \Rightarrow \frac{-i\eta}{p^2 + i\eta}$$

$$\Rightarrow \frac{-i\eta_{\mu\nu}}{p^2 + i\varepsilon}$$

$$\alpha$$

$$-ie\gamma^{\mu}_{\beta\alpha}(2\pi)^4\delta^{(4)}(p_1+p_2+p_3).$$

₩

$$-ie\gamma^{\mu}_{\beta\alpha}(2\pi)^4\delta^{(4)}(p_1+p_2+p_3).$$

#### Integrate over loop momenta

QED Feynman rules

$$\int \frac{d^a p}{(2\pi)^4}$$

Incoming fermion:

Incoming antifermion:





$$\implies u_{\alpha}(\mathbf{p},s)$$

Outgoing fermion:



 $\overline{u}_{\alpha}(\mathbf{p},s)$ 

 $\overline{v}_{\alpha}(\mathbf{p},s)$ 

Outgoing antifermion:



 $\nu_{\alpha}(\mathbf{p},s)$ 

Incoming photon:



 $\varepsilon_{\mu}(\mathbf{p})$ 

Outgoing photon:

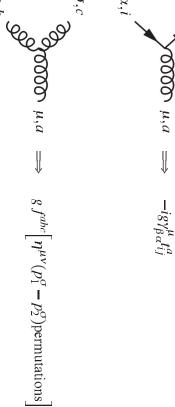


 $\varepsilon_{\mu}(\mathbf{p})^*$ 

every fermion loop and for every specially non-perturbatively. effort we can derive the Feynman rules A minus sign has to be included for bosons interactions, the are quite with the gluon or non-abelian vector electron positron ones, except that we quark factors are similar to the the next page. The quark and antifor QCD-like theories. They appear in to the final state. With some extra positron line that goes from the initial interesting physics perturbatively and involved and contain a large amount of numbers. The real difference comes need to include color quantum

## Standard Model Feynman rules

$$\alpha, i \longrightarrow \beta, j \implies \left(\frac{i}{\not p - m + i\varepsilon}\right)_{\beta\alpha} \delta_{ij}$$



$$\Rightarrow -ig^{2} \left[ f^{abe} f^{cde} \left( \eta^{\mu\sigma} \eta^{\nu\lambda} - \eta^{\mu\lambda} \eta^{\nu\sigma} \right) + \text{permutations} \right]$$

Although the rules seem to be those for QCD, notice that we could always include in the group theory factors t^a-{ij} chiral projectors and make the group not simple but semi-simple as in the case of the SM: SU(3)xSU(2)xU(1). If we work in nice renormalizable gauges, the only difference is that we have to include the Feynman rules for the couplings of the scalar sector. Something we will do later.

$$t^a_{ij} o t^a_{ij} \, rac{1}{2} (1 \pm \gamma_5)$$

With this simple trick the hard part, which is the coupling of the W,Z, and photons can be read simply from the rules in the LHS

## One example: Thomson Scattering

 $\gamma(k,\varepsilon) + e^{-}(p,s) \longrightarrow \gamma(k',\varepsilon') + e^{-}(p',s')$ 

simplicity but keeping explicitly the answer has to be a pure number times dependence on the photon polarisations. We can guess that the We work in the NR approximation for the classical electron radius



$$= (ie)^{2}\overline{u}(\mathbf{p}',s')\beta'(\mathbf{k}')^{*}\frac{\cancel{p}+\cancel{k}+m_{e}}{(p+k)^{2}-m_{e}^{2}}\beta(\mathbf{k})u(\mathbf{p},s)$$
$$+ (ie)^{2}\overline{u}(\mathbf{p}',s')\beta(\mathbf{k})\frac{\cancel{p}-\cancel{k}'+m_{e}}{(p-k')^{2}-m_{e}^{2}}\beta'(\mathbf{k}')^{*}u(\mathbf{p},s)$$

$$p^{2} = m_{e}^{2} = p'^{2}$$
$$k^{2} = 0 = k'^{2}$$

$$|\mathbf{p}|, |\mathbf{k}|, |\mathbf{p}'|, |\mathbf{k}'| \ll m_e$$
  $db = -bd + 2$ 

$$db = -bd + 2(a \cdot b)\mathbf{1}$$

$$(p+k)^2 - m_e^2 \approx 2m_e |\mathbf{k}|,$$

$$(p-k')^2 - m_e^2 \approx -2m_e |\mathbf{k'}|$$

$$(\not k - m)u(k,s) = 0.$$
  
 $\overline{u}(\mathbf{k},s)\gamma^{\mu}u(\mathbf{k},s) = 2k^{\mu}$ 

## Thomson Scattering, continued

$$\langle f|\widehat{S}|i\rangle = \langle f|i\rangle + (2\pi)^4 \delta^{(4)} \left(\sum_{\text{final}} p_i' - \sum_{\text{initial}} p_j\right) i \mathcal{M}_{i \to f}$$

$$d\sigma = \frac{|\mathcal{M}_{i \to f}|^2}{4E_1 E_2 |\mathbf{v}_1 - \mathbf{v}_2|} (2\pi)^4 \delta^{(4)} \left(p_1 + p_2 - \sum_{j=1}^n p_j'\right) d\Phi_k.$$

sum over the initial ones. We will consider unpolarised incoming degree of polarisation photons and study how the outgoing photons can gain some Square the amplitude, sum over final electron polarisations, and

$$F_{\text{coll}} = 4E_1 E_2 |\mathbf{v}_1 - \mathbf{v}_2| = 4E_1 E_2 \left| \frac{\mathbf{p}_1}{E_1} - \frac{\mathbf{p}_2}{E_2} \right|$$

$$= 4|E_2 \mathbf{p}_1 - E_1 \mathbf{p}_2| = 4\left(E_2 |\mathbf{p}_1| + E_1 |\mathbf{p}_2|\right)$$

$$= 4\sqrt{(p_1 \cdot p_2)^2 - m_1^2 m_2^2}.$$

$$\sum_{s=\pm 1} u_{\alpha}(\mathbf{k}, s) \overline{u}_{\beta}(\mathbf{k}, s) = (\not k + m)_{\alpha\beta}$$

$$\frac{d\sigma}{d\Omega} = \frac{1}{64\pi^2 m_e^2} |i\mathcal{M}_{i\to f}|^2 = \left(\frac{e^2}{4\pi m_e}\right)^2 |\varepsilon(\mathbf{k}) \cdot \varepsilon'(\mathbf{k}')^*|^2 \qquad \overline{d\Omega} = \overline{8}$$

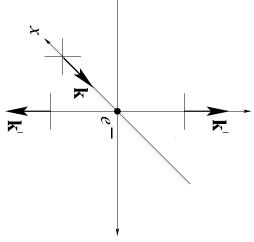
$$\frac{d\sigma}{d\Omega} = \frac{3}{8\pi} \sigma_T \left| \varepsilon(\mathbf{k}) \cdot \varepsilon'(\mathbf{k}')^* \right|^2$$

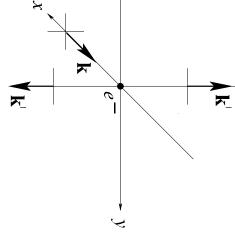
$$\frac{4}{m_e^2} = \frac{8\pi}{3} r_{\text{cl}}^2 \qquad \frac{1}{2} \sum_{a=1,2} \left| \varepsilon(\mathbf{k}, a) \cdot \varepsilon'(\mathbf{k}')^* \right|^2 = \frac{1}{2} \left( \delta_{ij} - \frac{k_i k_j}{|\mathbf{k}|^2} \right) \varepsilon'_j(\mathbf{k}') \varepsilon'_i(\mathbf{k}')^* \\
= \frac{1}{2} \left[ 1 - |\hat{\mathbf{k}} \cdot \varepsilon'(\mathbf{k}')|^2 \right],$$

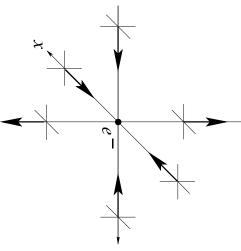
We want to monitor the polarisation of the outgoing photons even when the incoming ones are not polarised

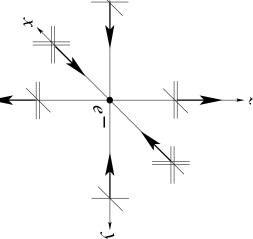


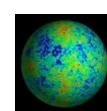
## Thomson and CMB Polarisation









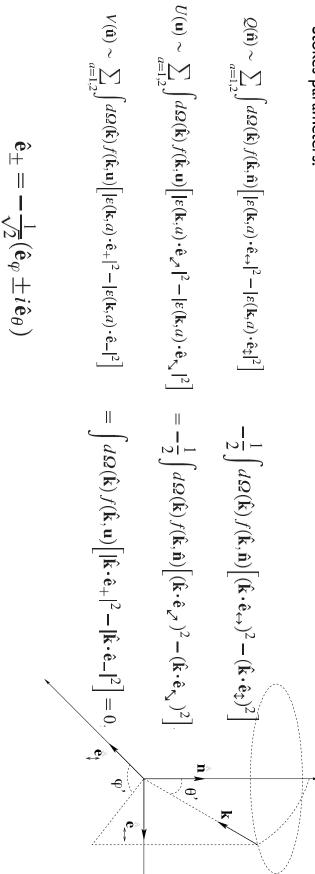


How we can get polarised light

Stokes parameters:

light does not generate polarisation

An isotropic incoming distribution of A incoming light with a quadrupole perturbation generates net polarisation



## Quadrupole distribution

Finally we reach the punch line. No circular polarisation is generated by Thomson scattering, and we can write the combination:

$$Q(\hat{\mathbf{n}}) \pm iU(\hat{\mathbf{n}}) \sim -\int d\Omega(\theta', \varphi') f(\theta', \varphi'; \hat{\mathbf{n}}) \sin^2 \theta' e^{\pm 2i\varphi'}$$

$$Y_2^{\pm 2}(\theta', \varphi') = 3\sqrt{\frac{5}{96\pi}} \sin^2 \theta' e^{\pm 2i\varphi'}$$

gravitational waves. Inflation predicts primordial gravitation waves, the has unexpected consequences obtain this information. The simple computation of Thomson scattering measurement of polarisation in the CMB offers an amazing window to One of the obvious generators of quadrupole anisotropies are

$$Q(\hat{\mathbf{n}}) \pm iU(\hat{\mathbf{n}}) = -\sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \left( E_{\ell m} \pm iB_{\ell m} \right) \pm 2Y_{\ell}^{m}(\hat{\mathbf{n}})$$

$$\langle E_{\ell m}^* E_{\ell' m'} \rangle = C_{\ell}^{EE} \delta_{\ell \ell'} \delta_{mm'}, \qquad \langle B_{\ell m}^* B_{\ell' m'} \rangle = C_{\ell}^{BB} \delta_{\ell \ell'} \delta_{mm'}$$



#### Noether's Theorem

or anti-unitary preserve probability amplitudes (their modulus). They can be unitary In a QM theory physical symmetries are maps among states that Quantum mechanical realisation of Symmetries (Wigner's theorem).



$$|\alpha\rangle \longrightarrow |\alpha'\rangle$$

$$\beta\rangle \longrightarrow |\beta'\rangle$$

$$|\langle \alpha | \beta \rangle| = |\langle \alpha' | \beta' \rangle|.$$

$$\langle \mathcal{U} \alpha | \mathcal{U} \beta \rangle = \langle \alpha | \beta \rangle$$

$$\langle \mathscr{U} \alpha | \mathscr{U} \beta \rangle = \langle \alpha | \beta \rangle^*$$

**EQUATIONS OF MOTION you can show that:** under infinitesimal transformations, AND WITHOUT USING THE For continuous symmetries we have Noether's celebrated theorem: If

 $\delta_{\varepsilon} \mathcal{L} = \partial_{\mu} K^{\mu}$ 

$$S[\phi] = \int d^4 x \, {\cal L}(\phi, \partial_\mu \phi)$$

then there is a conserved current in the theory

#### Noether's Theorem

#### In formulas:

$$\begin{split} \delta_{\varepsilon}\mathcal{L} &= \frac{\partial \mathcal{L}}{\partial (\partial_{\mu}\phi)} \partial_{\mu} \delta_{\varepsilon} \phi + \frac{\partial \mathcal{L}}{\partial \phi} \delta_{\varepsilon} \phi \\ &= \partial_{\mu} \left( \frac{\partial \mathcal{L}}{\partial (\partial_{\mu}\phi)} \delta_{\varepsilon} \phi \right) + \left[ \frac{\partial \mathcal{L}}{\partial \phi} - \partial_{\mu} \left( \frac{\partial \mathcal{L}}{\partial (\partial_{\mu}\phi)} \right) \right] \delta_{\varepsilon} \phi \\ &= \partial_{\mu} K^{\mu}. \end{split}$$

$$\partial_{\mu}J^{\mu}=0$$

$$J^{\mu} = \frac{\partial \mathcal{L}}{\partial (\partial_{\mu} \phi)} \delta_{\varepsilon} \phi - K^{\mu}$$

With a conserved charge that generates the symmetry:

$$Q \equiv \int d^3x J^0(t, \mathbf{x})$$

$$\frac{dQ}{dt} = \int d^3x \, \partial_0 J^0(t, \mathbf{x}) = -\int d^3x \, \partial_i J^i(t, \mathbf{x}) = 0,$$

$$\delta \phi = i[\phi, Q].$$

Space-time translations -- Energy-Momentum Lorentz transformation -- Angular momentum and CM motion Phase rotation -- abelian and non-abelian charges

Massive Dirac fermions:

$$\mathcal{L} = i \overline{\psi}_j \beta \psi_j - m \overline{\psi}_j \psi_j$$

$$\psi_i \longrightarrow U_{ij}\psi_j$$

 $U \in U(N)$  Nthe number of fermions

Useful examples

$$U = \exp(i\alpha^a T^a),$$

$$(T^a)^{\dagger} = T^a$$

$$j^{\mu a} = \overline{\psi}_i T^a_{ij} \gamma^\mu \psi_j$$

$$\partial_{\mu}j^{\mu}=0$$

$$Q^a = \int d^3x \psi_i^{\dagger} T_{ij}^a \psi_j$$

$$[Q^a, H] = 0.$$

$$\mathscr{U}(\alpha) = e^{i\alpha^a \mathcal{Q}^a}$$

When U is the identity, we have fermion number, or charge

In the m=0 we have more symmetry: CHIRAL SYMMETRY, rotate L,R fermions independently

$$\mathcal{L} = i ar{\psi}_{jL} 
ot \partial_{} \psi_{Lj} + i ar{\psi}_{jR} 
ot \partial_{} \psi_{Rj}$$
  $\psi_{\mathbb{L},\mathbb{R}} o U_{\mathbb{L},\mathbb{R}} \psi_{\mathbb{L},\mathbb{R}} \quad U(N)_{\mathbb{L}} imes U(N)_{\mathbb{R}}$